

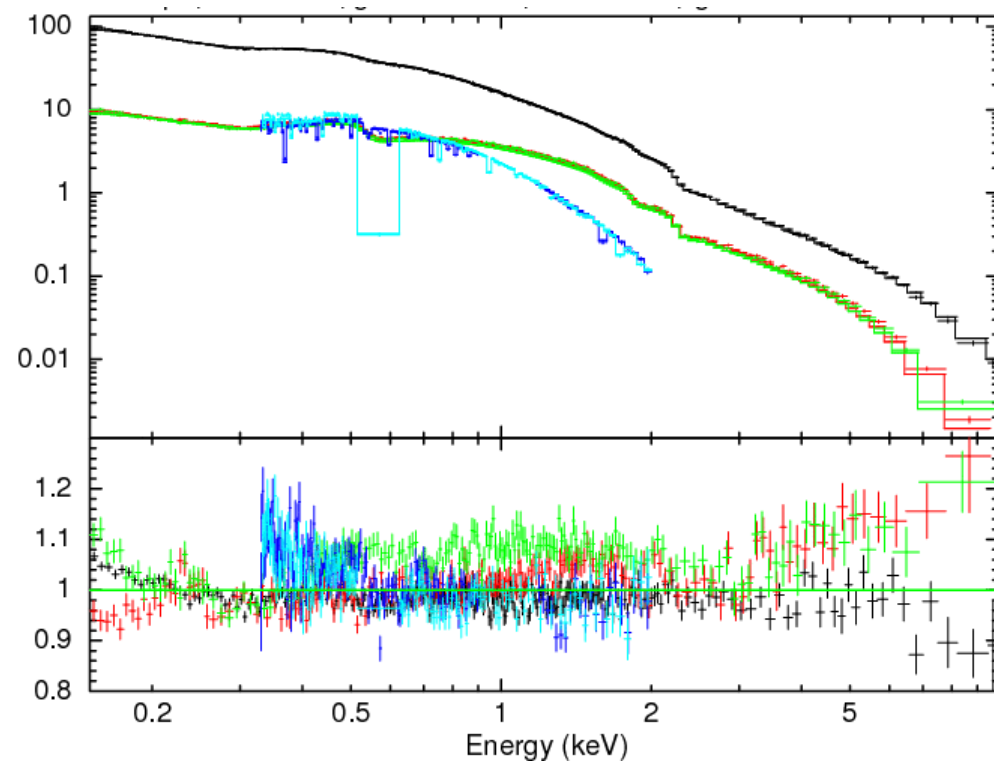
XRT PSF Investigations

Michael Smith, ESAC

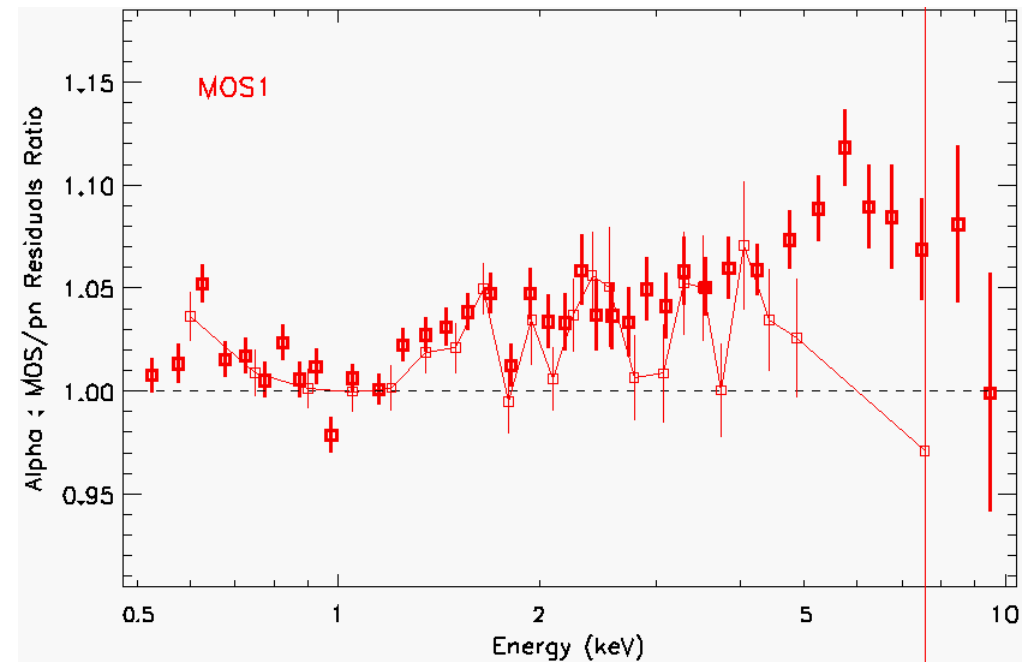
EPIC Cal Ops Meeting, MPE, 27-28 April 2017

PKS 2155-304

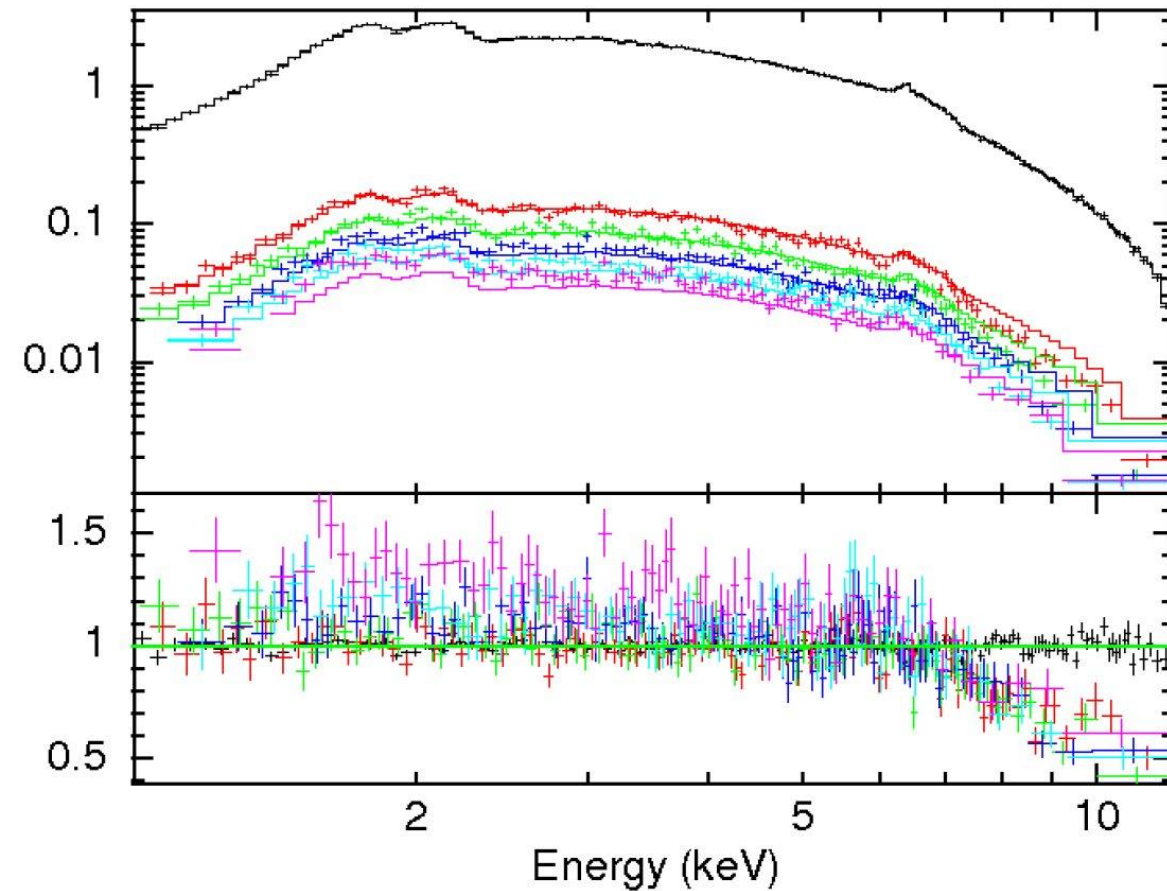
- Piled-up
- Annular extraction regions for EPIC



Sample of non-piled-up on-axis sources



Read et al. 2014



NGC 5506 (Seyfert 2)

Dominated (>90%) by point source emission.

PN spectral comparison:

- annuli from 20''-25'' to 40''-45''
- compared with 20'' circle

Similar results for HE1136-2304 (in out-burst)

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Hopefully can be improved by tuning the PSF model parameters

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Current ELLBETA implementation: 240 parameters per XRT

- 5 PSF model parameters (r_0 , α , ϵ , FWHM, norm)
- 8 energies (0.10, 1.50, 2.75, 4.25, 6.00, 8.00, 10.25, 15.00 keV)
- 6 off-axis angles (0', 3', ..., 15')

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Limit the problem to 14 parameters per XRT

- 2 PSF model parameters (r_0 , α)
- 7 energies (15.00 keV \equiv 10.25 keV)
- 1 off-axis angle ($0'$ for MOS1, $1.2'$ for MOS2 and PN)

Iterative scheme for the empirical correction of XRT XPSF parameters

- ❖ Per source:
 - Extract spectrum from circular region
 - Extract spectra from several annular regions
 - Create respective RMFs
 - Create respective ARFs
 - Fit model to circular region spectrum
 - Apply this reference model to the annular spectra

- ❖ For all annular spectra and all sources, determine a suitable statistic, e.g.:

$$\sum_i (d_i - m_i)^2 / e_i$$

- ❖ Modify ELLBETA parameters in order to minimise the statistic

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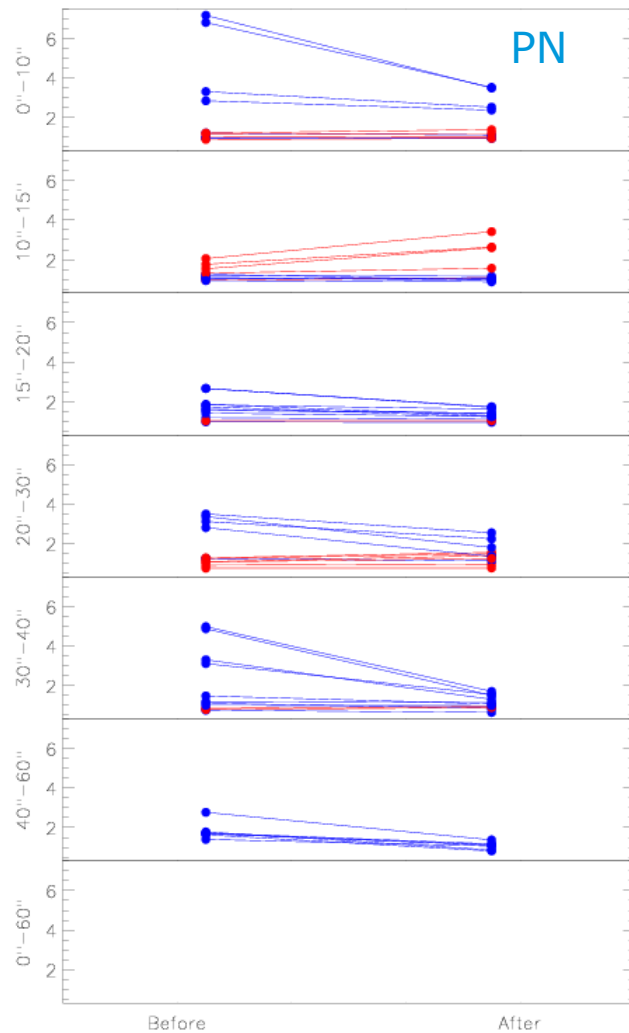
Source sample consists of 11 observations of bright non-piled-up point sources located at the nominal aim point:

		PN	MOS1	MOS2
MS0205.7+3509	0084140101	FF	FF	FF
Holmberg IX	0112521001	FF	FF	FF
Markarian 6	0144230101	FF	FF	FF
PKS B1334-127	0147670201	FF	FF	FF
RX J0228-40	0200480101	FF	FF	FF
LBQS 1228+1116	0306630201	FF	FF	FF
NGC 5506	0554170101	SW	--	--
NGC 5506	0554170201	SW	--	--
RBS 1055	0555020201	FF	FF	FF
MCG-5-23-16	0727960101	SW	--	--
MCG-5-23-16	0727960201	SW	--	--

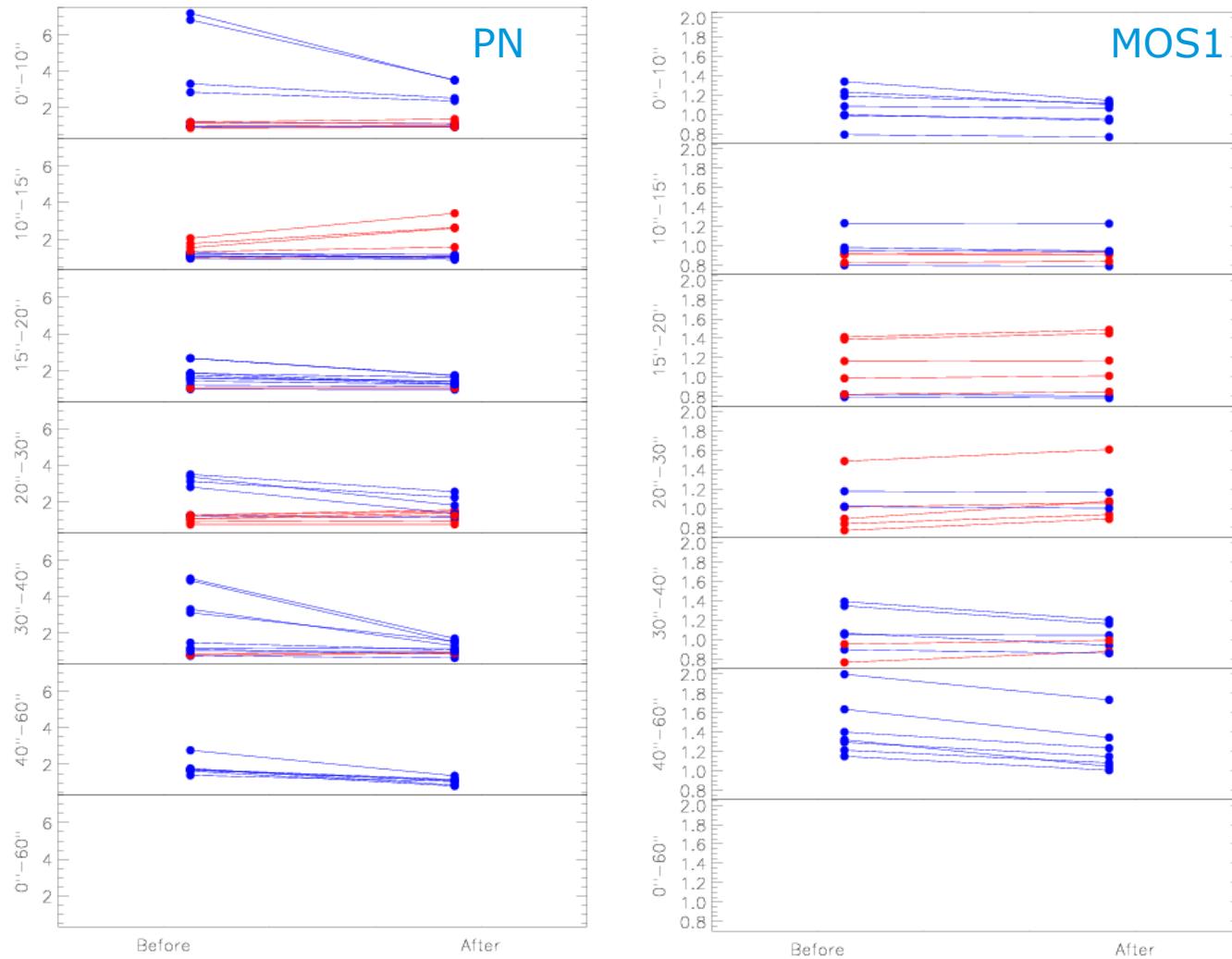
5-6 nested annuli being compared (up to 60" outer radius)

2 PSF model parameters (r_0 and α) being varied

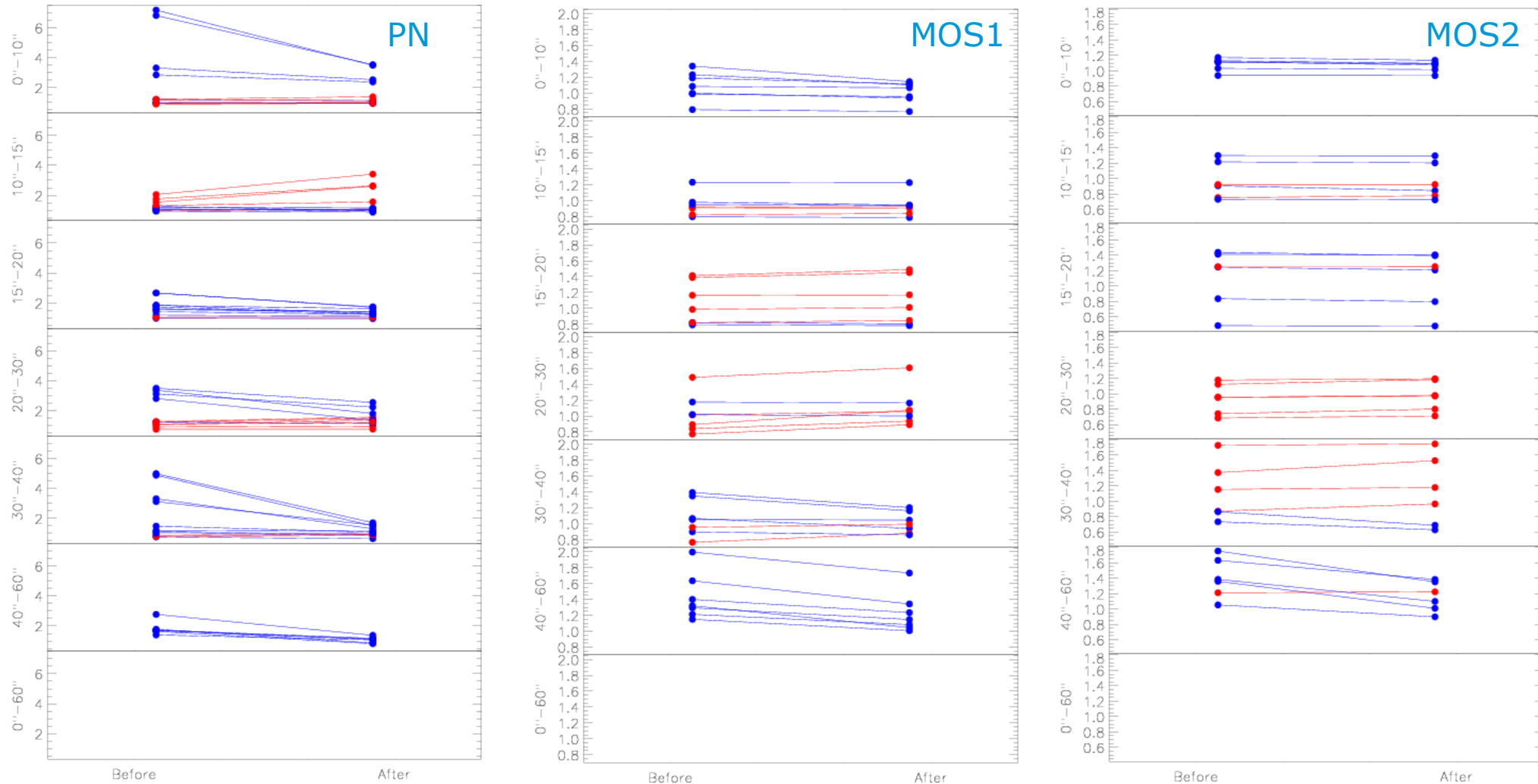
Change in Chi2 from public to candidate CCFs



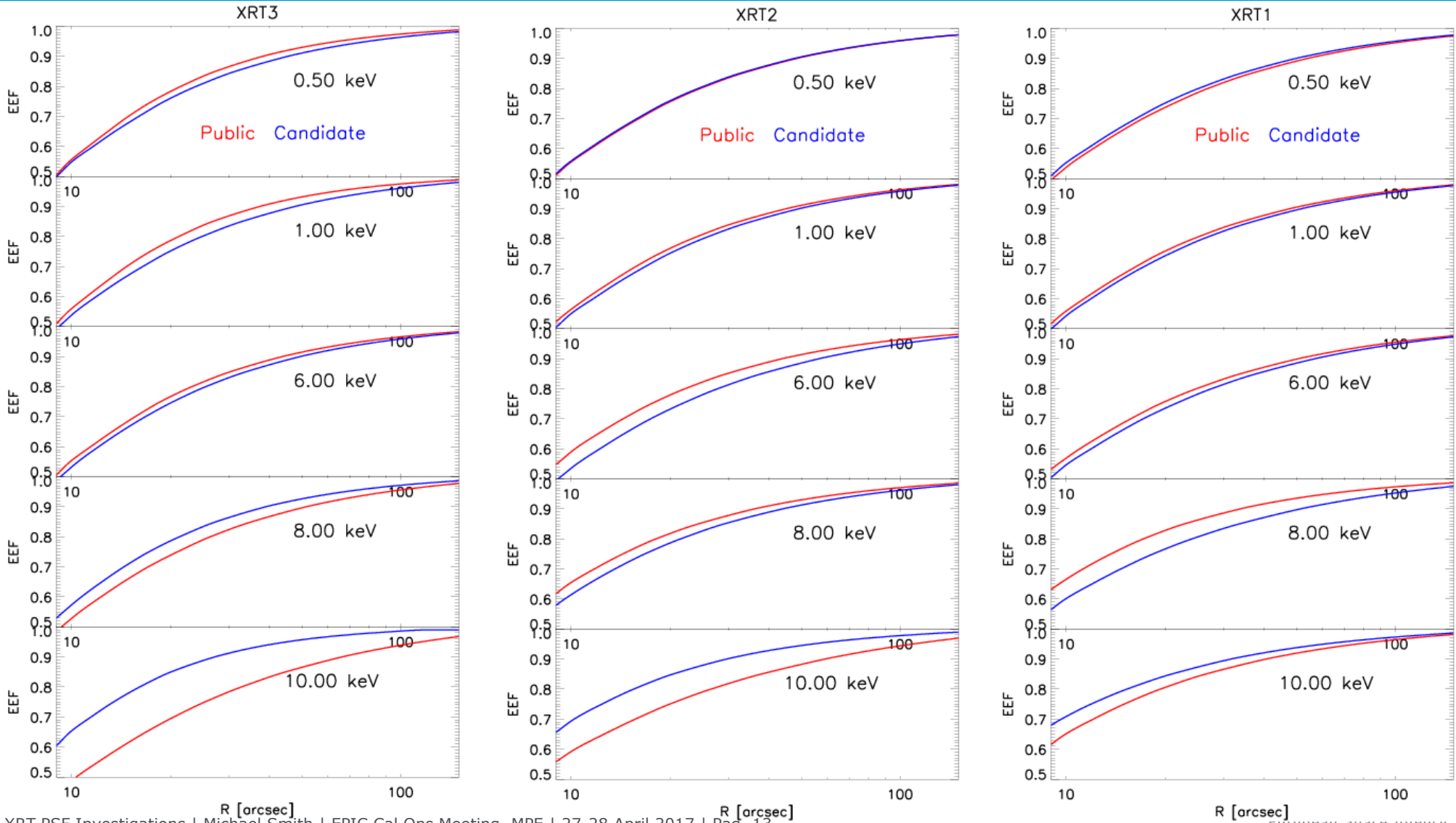
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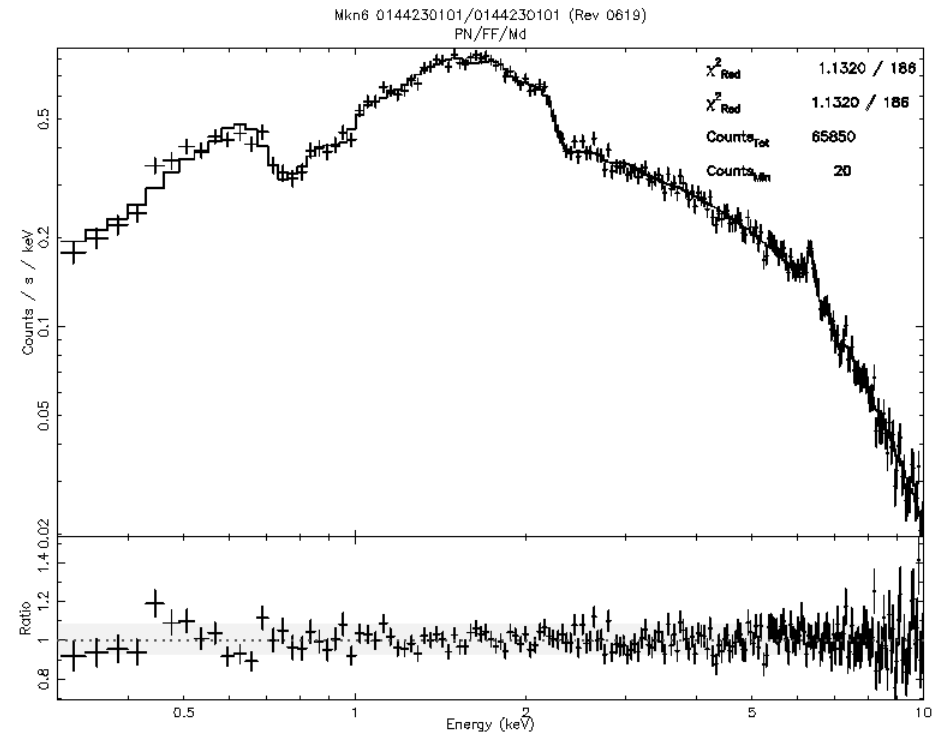
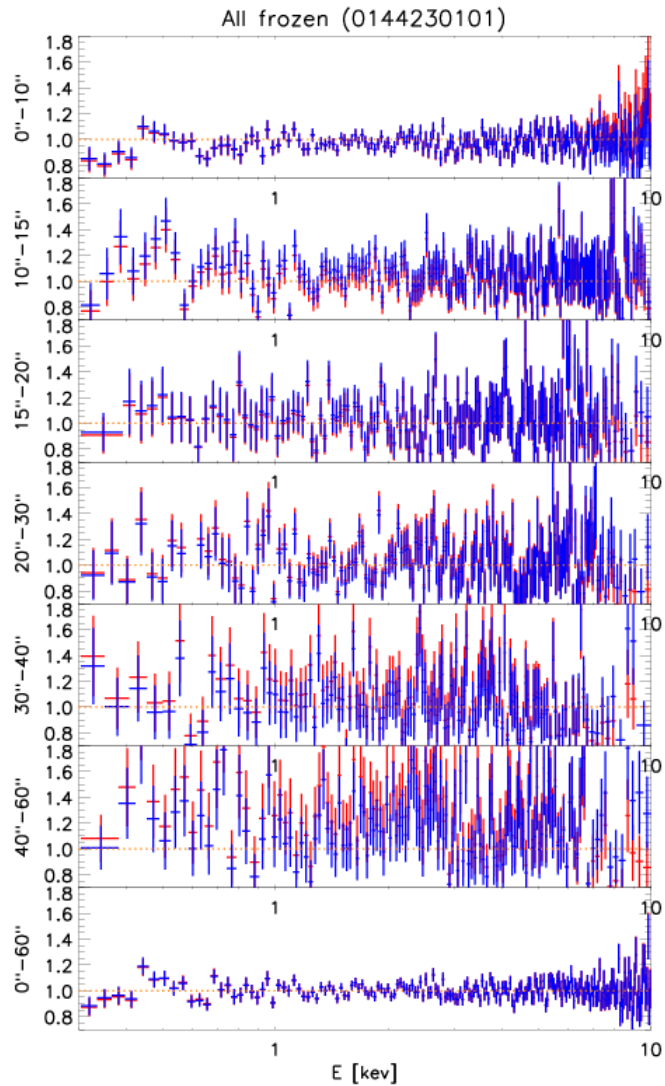
Change in Chi2 from public to candidate CCFs



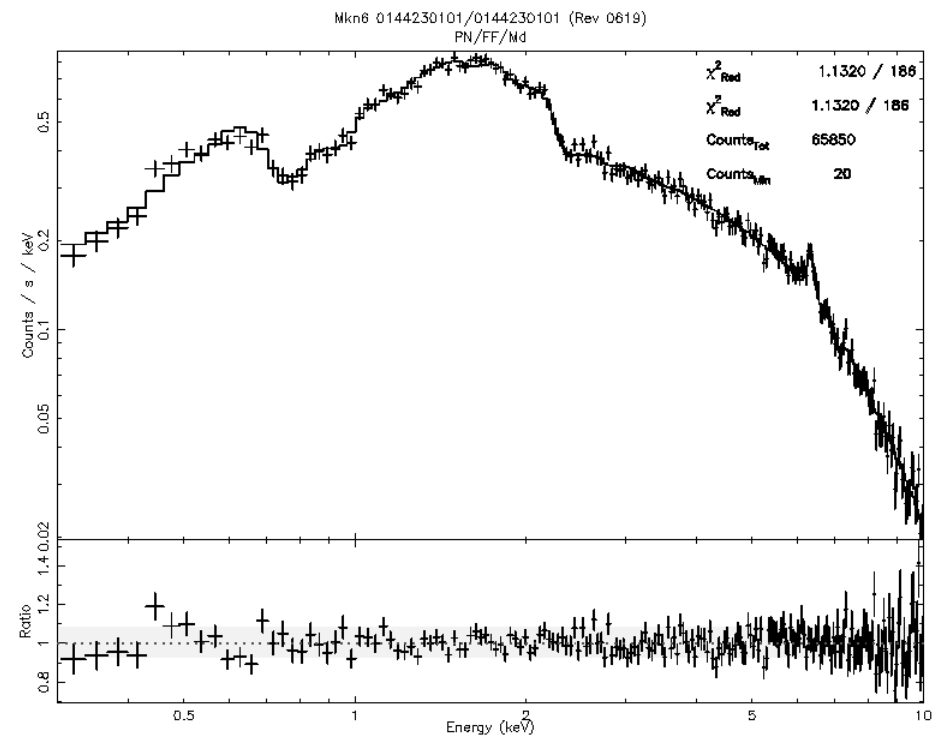
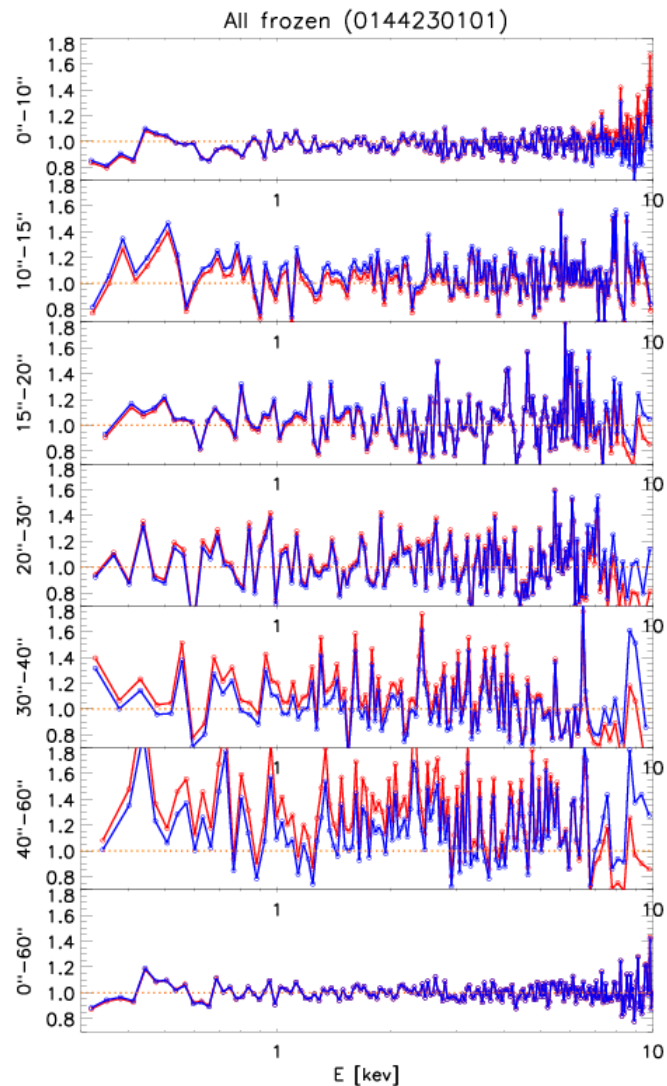
Comparison of EEFs



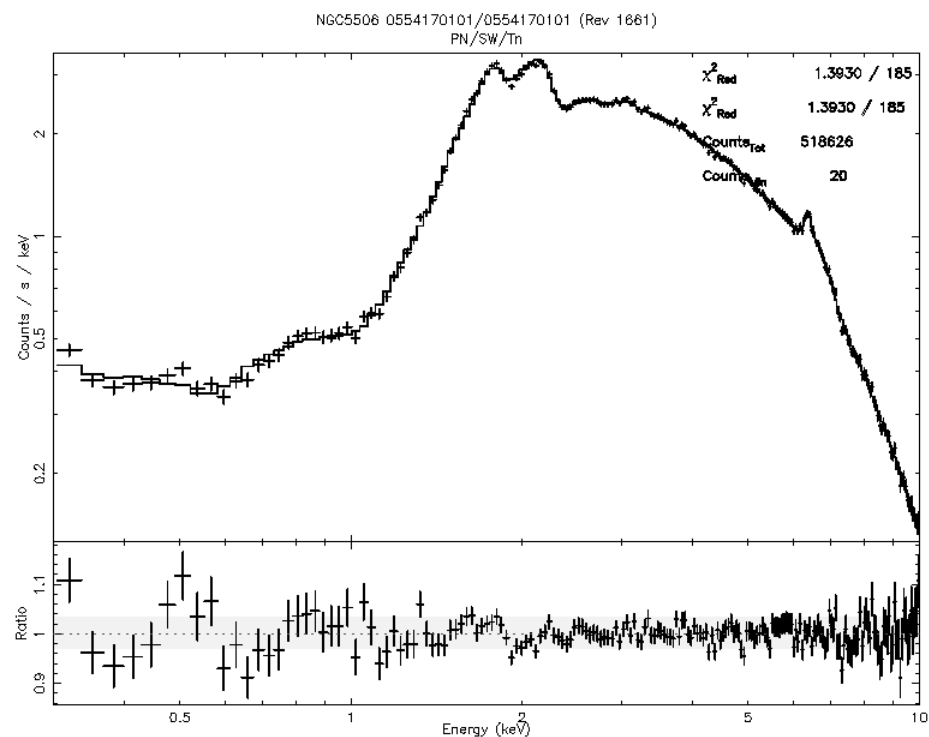
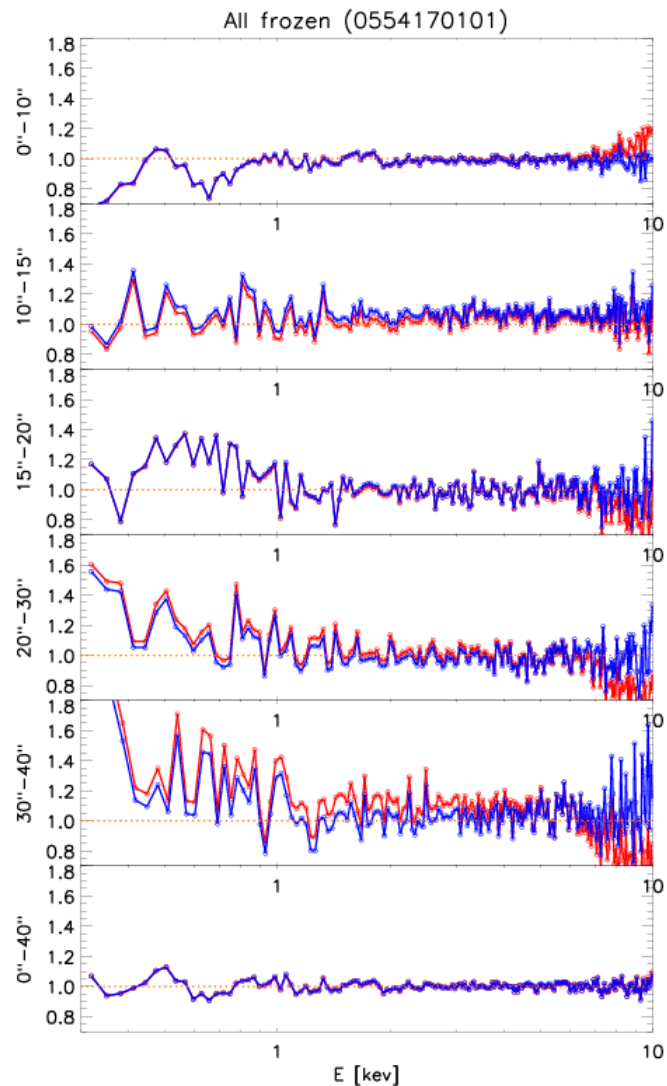
Results: Mkn 6 (PN)



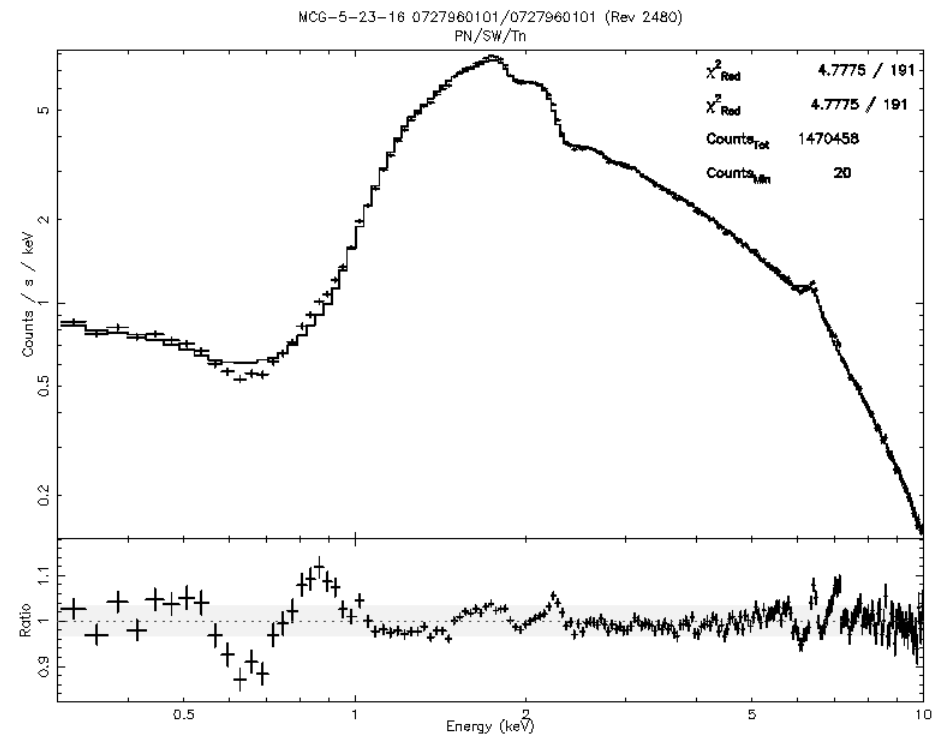
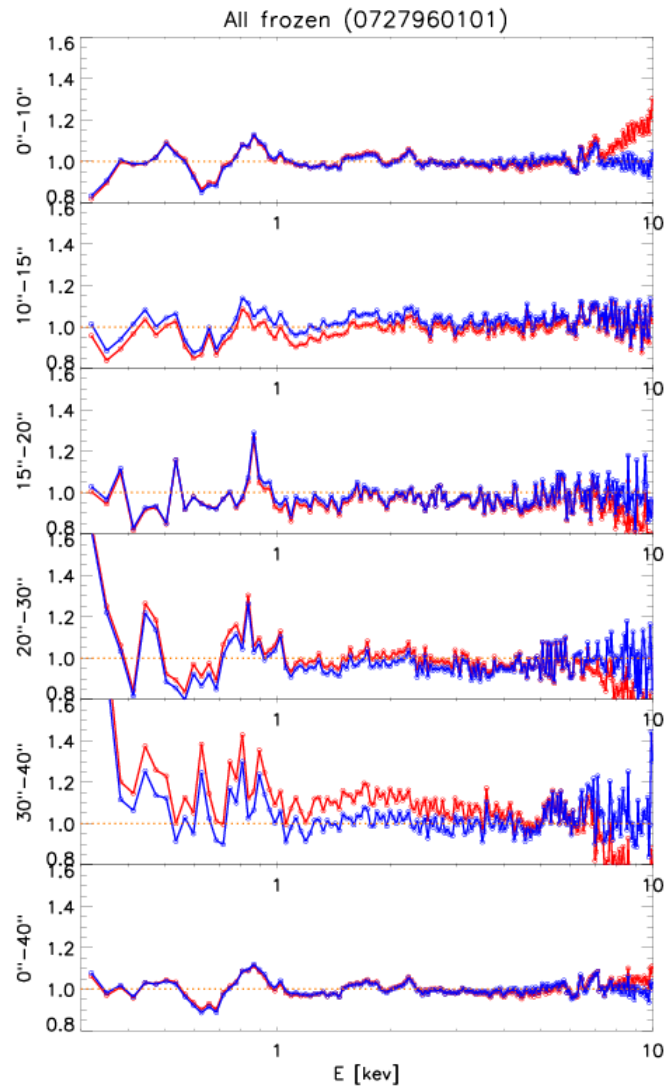
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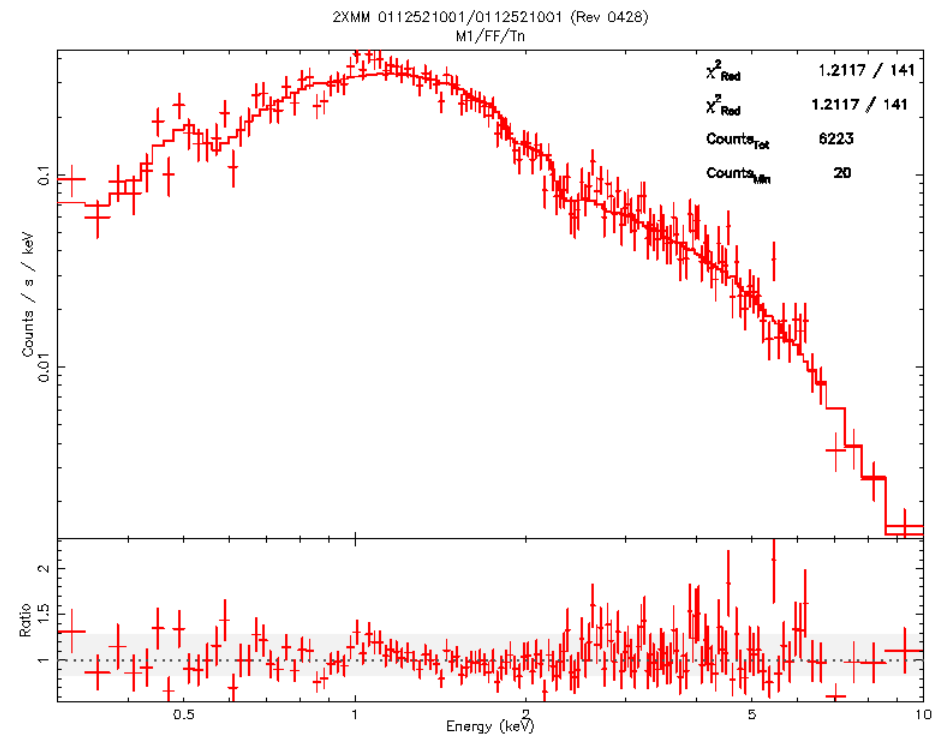
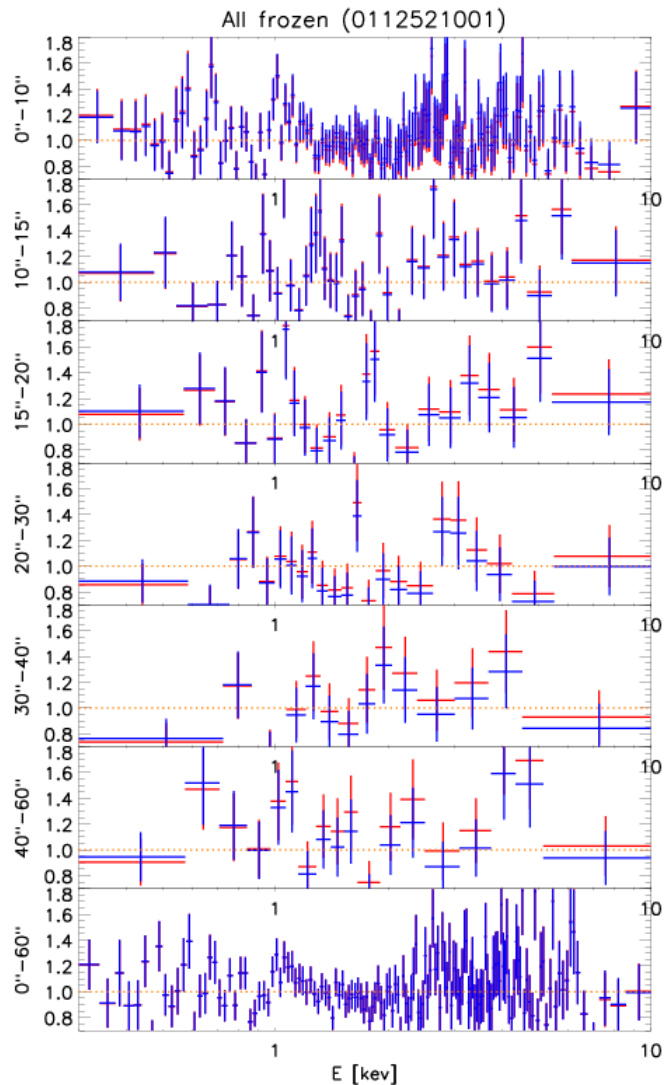
Results: NGC 5506 (PN)



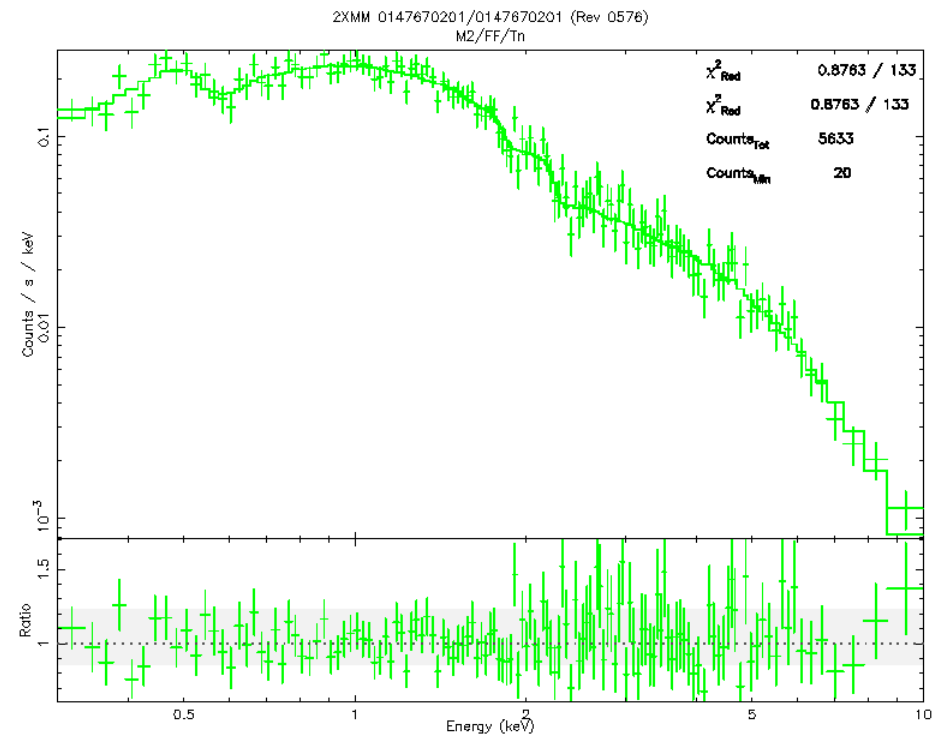
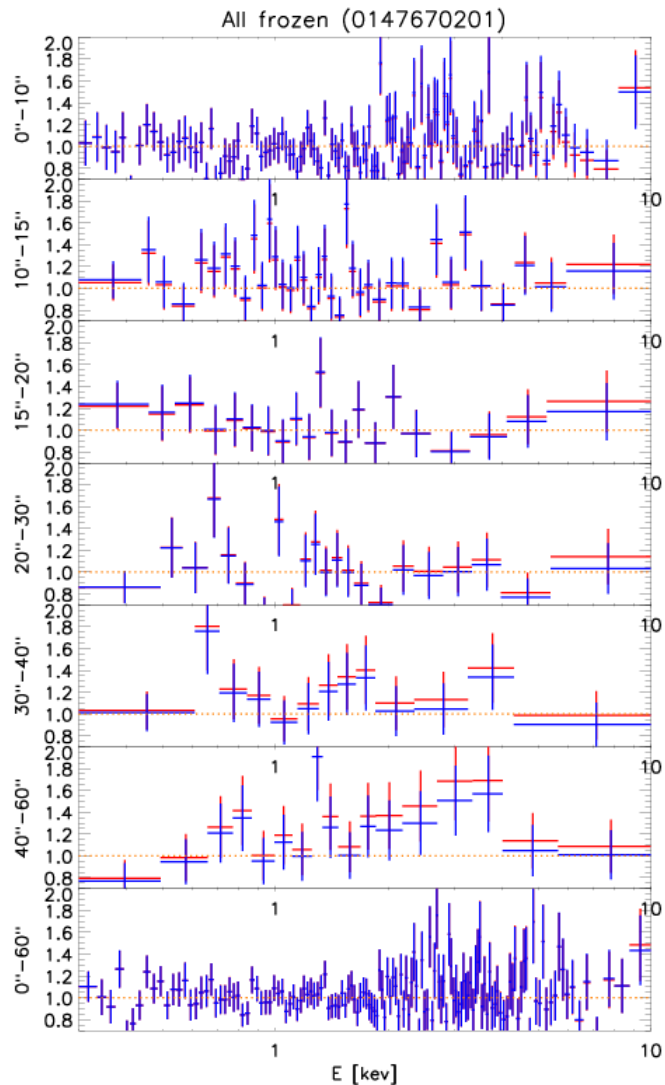
Results: MCG-5-23-18 (PN)



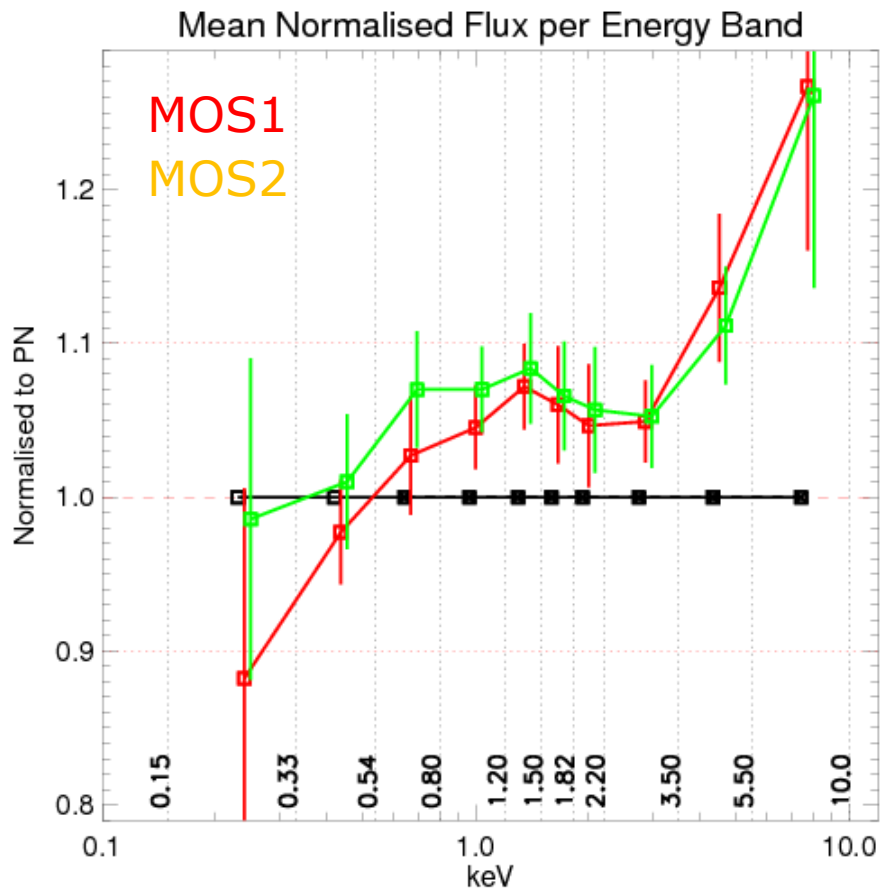
Results: Holmberg IX (MOS1)



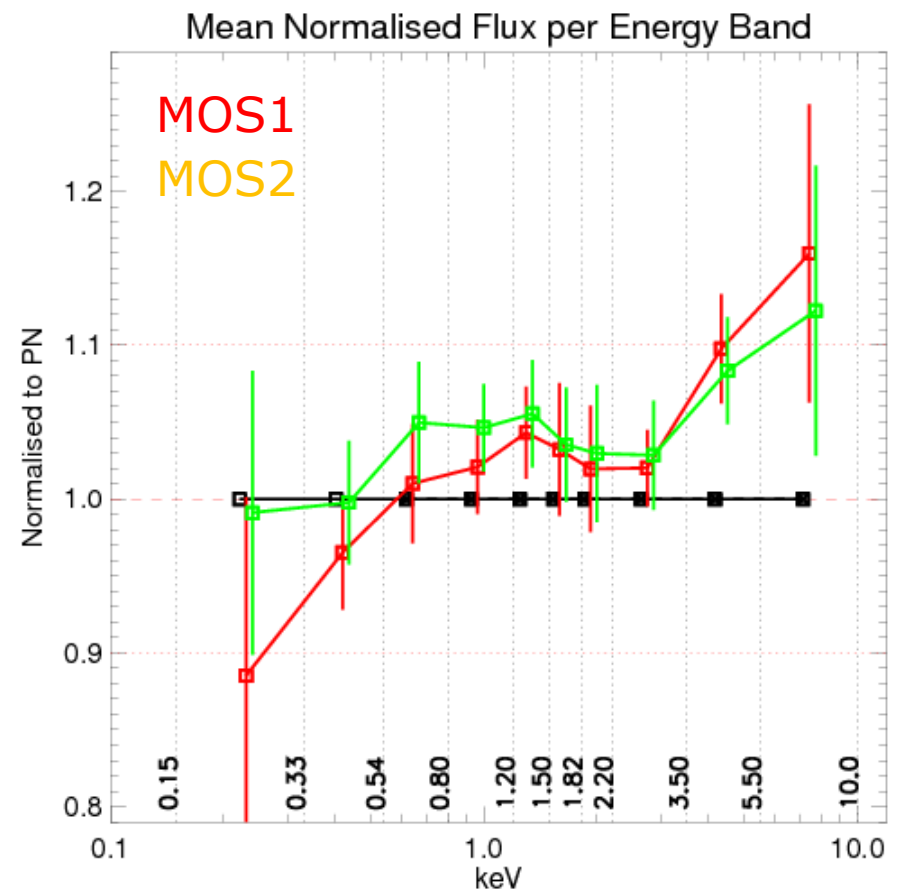
Results: PKS B1334-127 (MOS2)



Public XRTn CCFs



Candidate XRTn CCFs



Further validation:

- Confirm no drastic change in overall spectra
- Check > 10 keV
- Confirm no impact on source detection

