

X-RAY EMISSION OF INTERMEDIATE-MASS T TAURI STARS

María Alejandra Díaz¹, Gwendolyn Meeus¹, Jari Kajava², Ignacio Mendigutía³

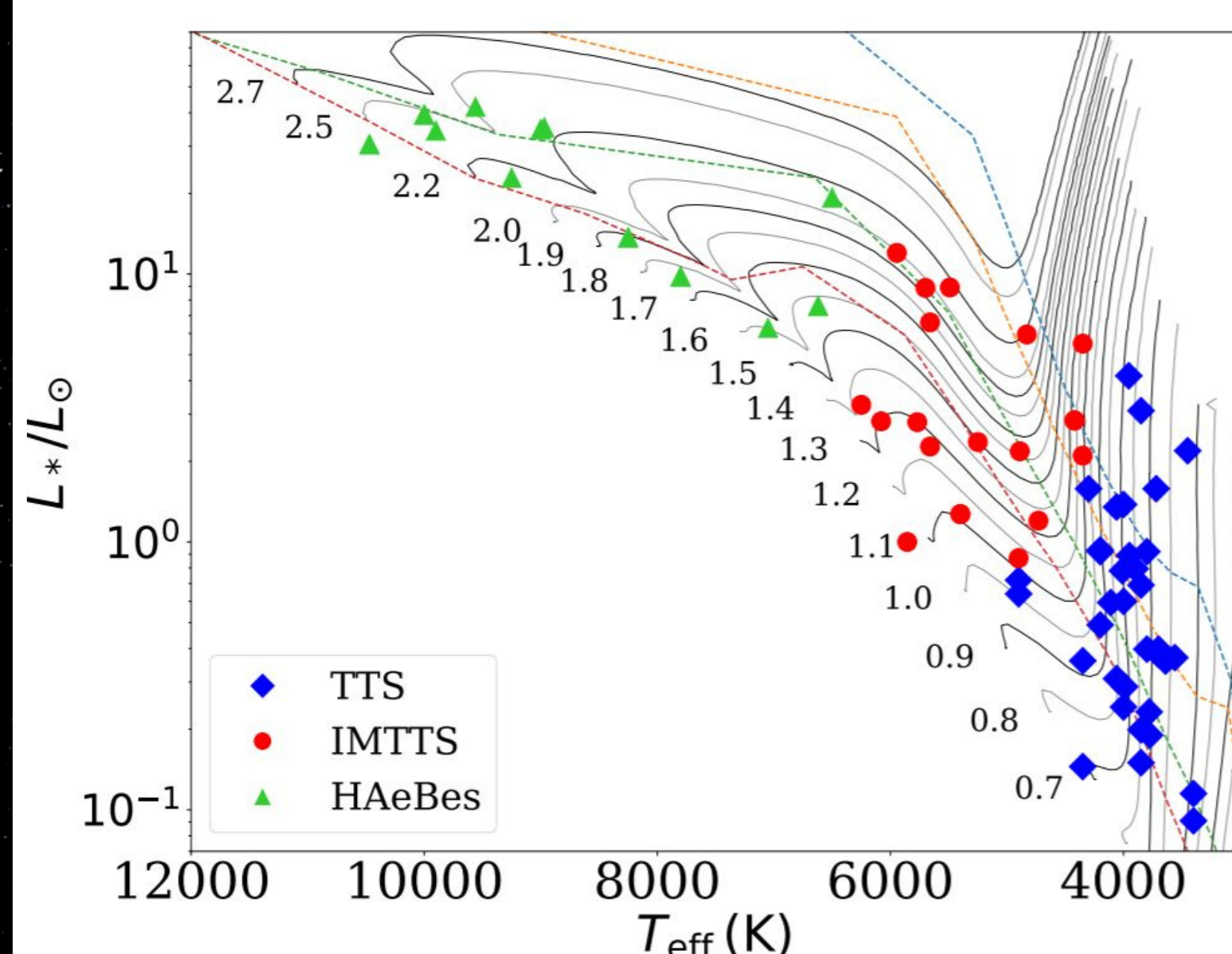


What are IMTTS?

Intermediate-mass T Tauri stars (IMTTS) are pre-main sequence (PMS) stars surrounded by protoplanetary disks. They are more massive than the solar-mass **T Tauri stars** (TTS), and younger than the intermediate-mass **Herbig Ae/Be stars** (HAEBEs).

Why are they interesting?

In stars, convection + rotation can generate a magnetic dynamo.



Most PMS stars start out fully **convective**, but while the low-mass **TTS** take a long time to develop a **radiative** core (some even never do), **HAEBs** do so rather fast.

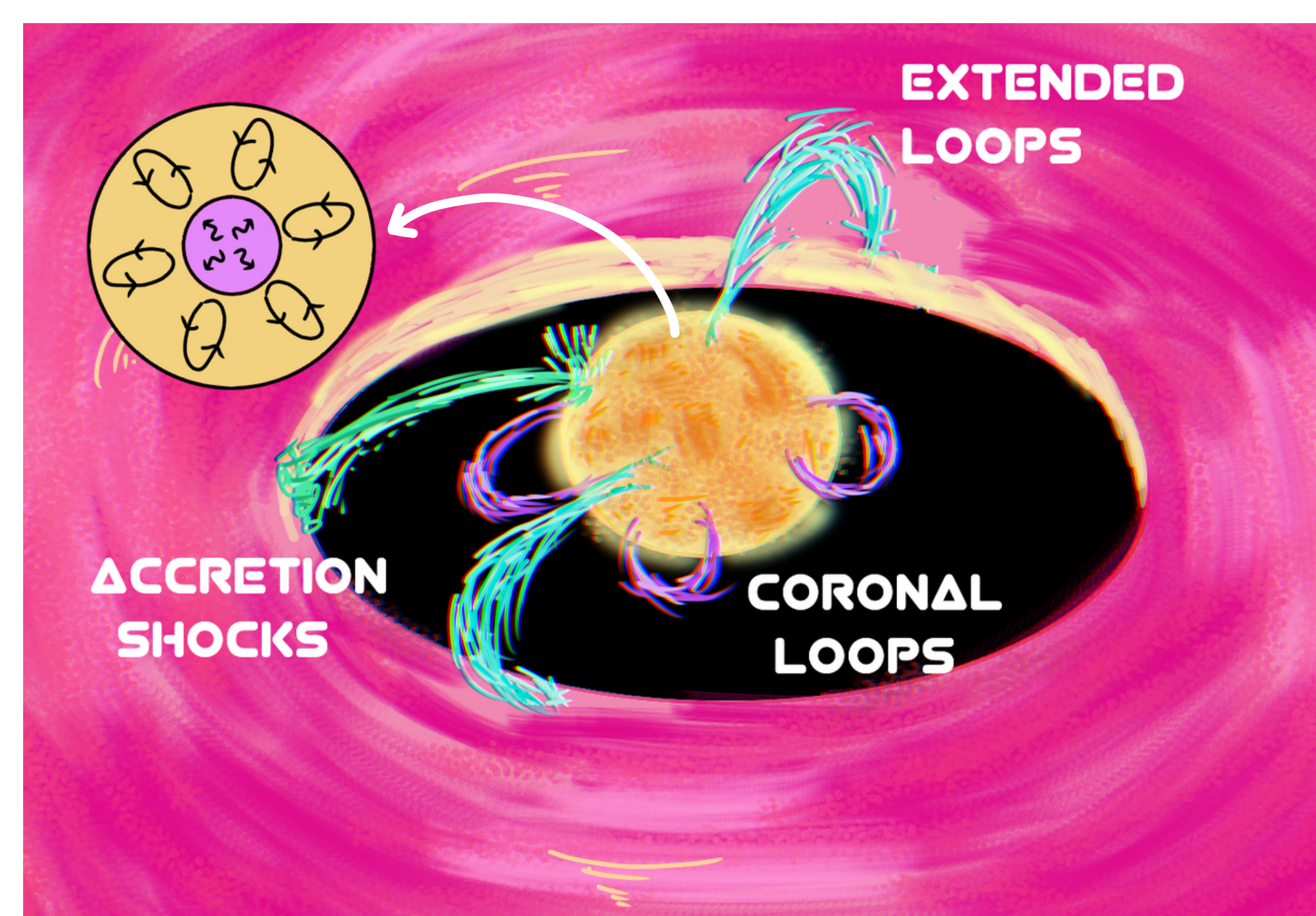
Although **HAEBs** are expected to lack the magnetic fields necessary to have **coronal X-ray emission**, they have been detected in X-rays. The reason for this is still debated, which is why we want to study their precursors, the **IMTTS**. Possible scenarios include a hidden **T Tauri** companion or remnant magnetic fields.

Aim of the study

To study the **X-ray properties** of **IMTTS** to shed light onto the magnetic evolution of intermediate PMS stars.

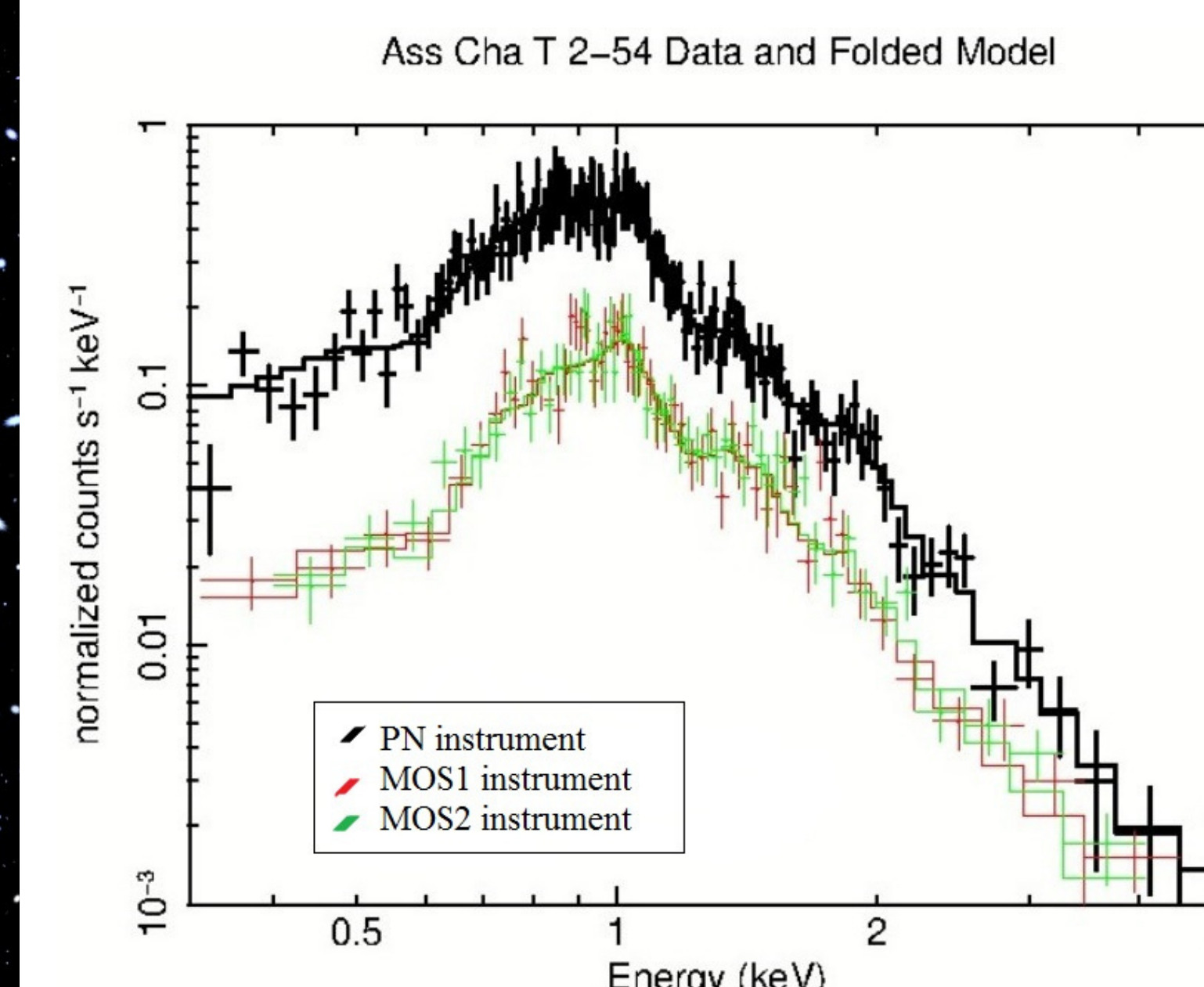
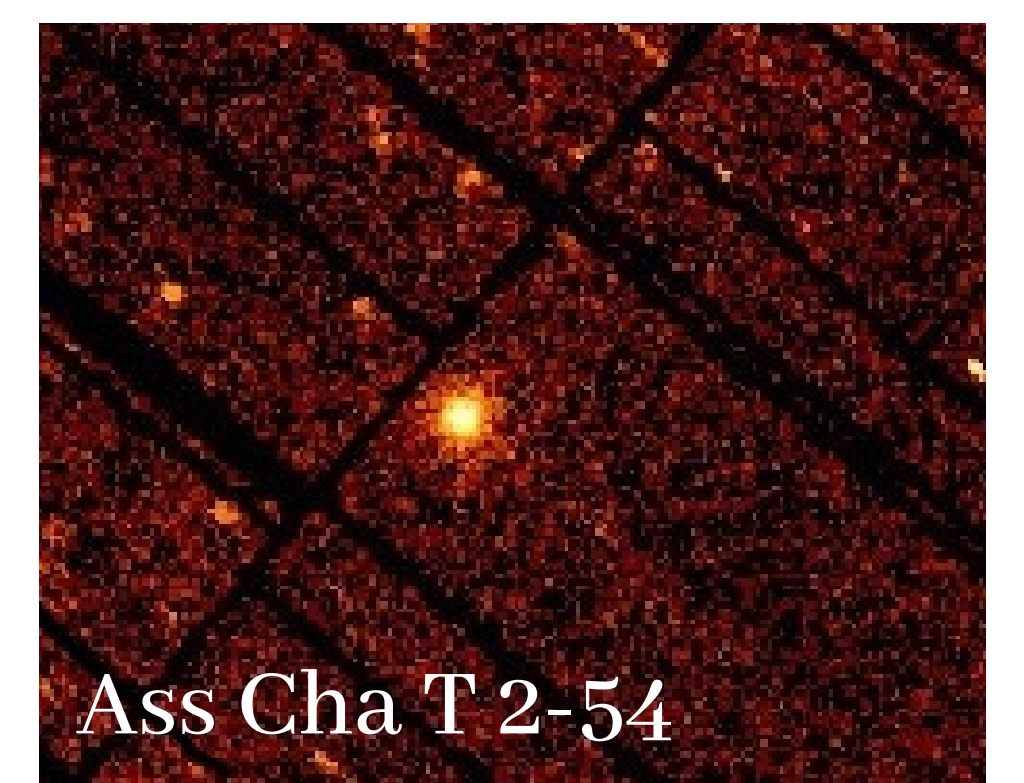
How do we do this?

- We compiled a sample of 60+ **IMTTS**
- We searched for archival data and obtained **new observations** with XMM
- We are currently analyzing the new X-ray data
- We will relate the **X-ray properties** with stellar and disc properties, such as **radiative** core radius and accretion.
- Finally, we will compare the results with those of **HAEBs** and **T Tauri** stars



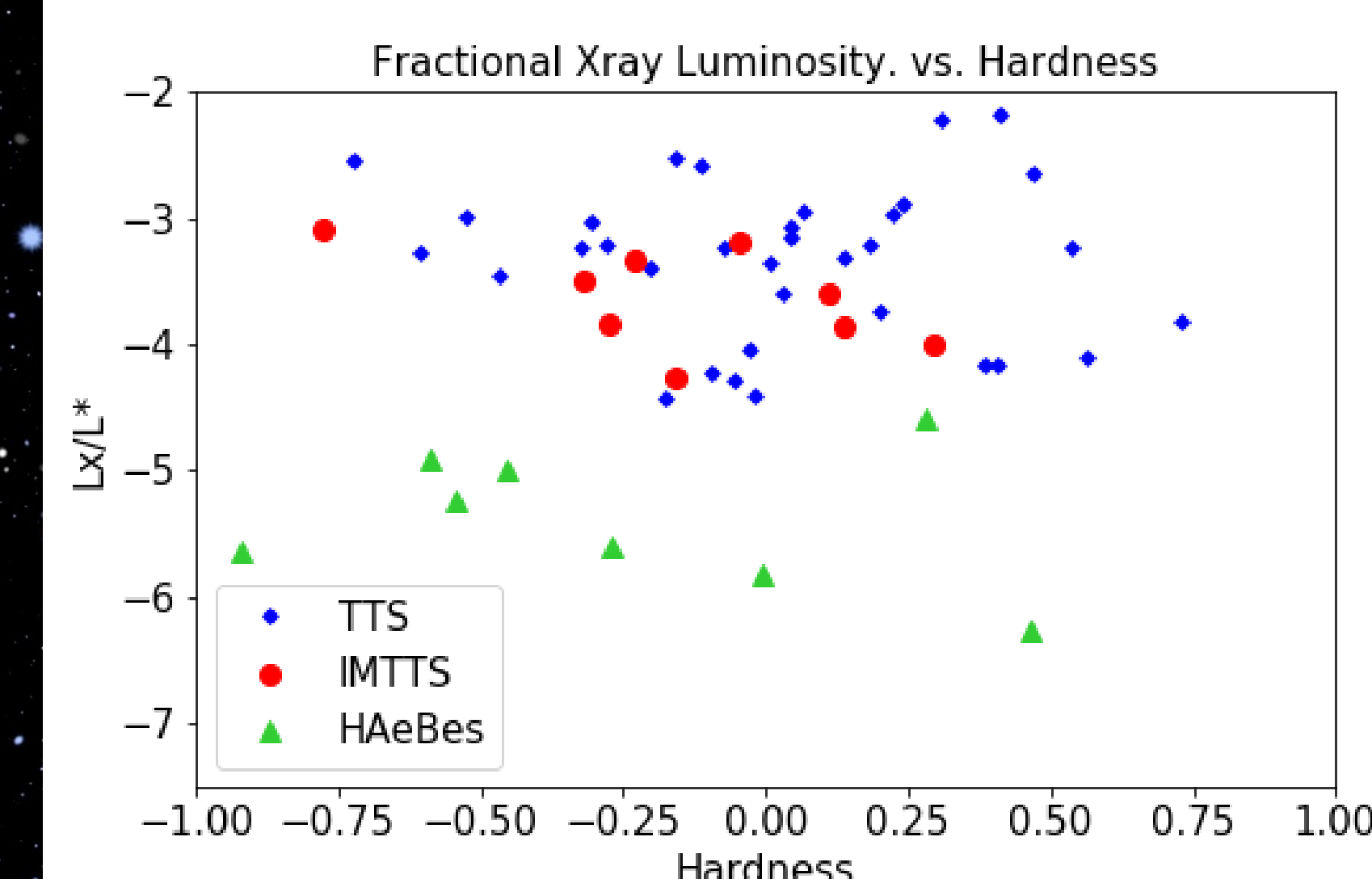
Preliminary results

We obtained observations for 9 **IMTTS** that had not yet been observed in **X-rays**. Most objects emit strongly in X-rays.



X-ray spectra of **IMTTS** are well described with two-component APEC models, as is also observed in many **TTS** and **HAEBs**.

While the hardness ratio has a similar range, the fractional X-ray luminosity in **HAEBs** is lower than that of **IMTTS** and **TTS**. This might be related to the later development of a radiative core.



Next steps

More detailed analysis for the spectra, also those from the archive.



I am looking for a PhD position! If you want to talk to me, you can contact me at madiazteo@gmail.com