

Unveiling stellar wind structures in HMXBs: *A high-resolution study of Vela X-1 with XMM-Newton*

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An example: Vela X-1

(1)	Distance	$2.42^{+0.19}_{-0.16}$ kpc
(2)	HD 77581 radius R_{\star}	$31^{+4}_{-3} R_{\odot}$
(3)	HD 77581 mass M_{\star}	$21.5 \pm 4.0 M_{\odot}$
(4)	Neutron star mass M_{NS}	$\sim 1.8 M_{\odot}$
(5)	Orbital period P_{orb}	8.964357 ± 0.000029 day
(6)	Semi-major axis $a \sin i$	113.89 ± 0.13 lt-sec
(6)	Eccentricity e	0.0898 ± 0.0012
(6)	Longitude of periastron ω	152.59 ± 0.92
(7)	Inclination i	$> 73^{\circ}$
(8)	Pulse period	~ 283 sec
(7)	Orbital separation a	$\sim 1.8 R_{\star}$

(1) *Bailer-Jones et al. (2018)*

(2) *Joss & Rappaport (1984)*

(3) *Giménez-García et al. (2016)*

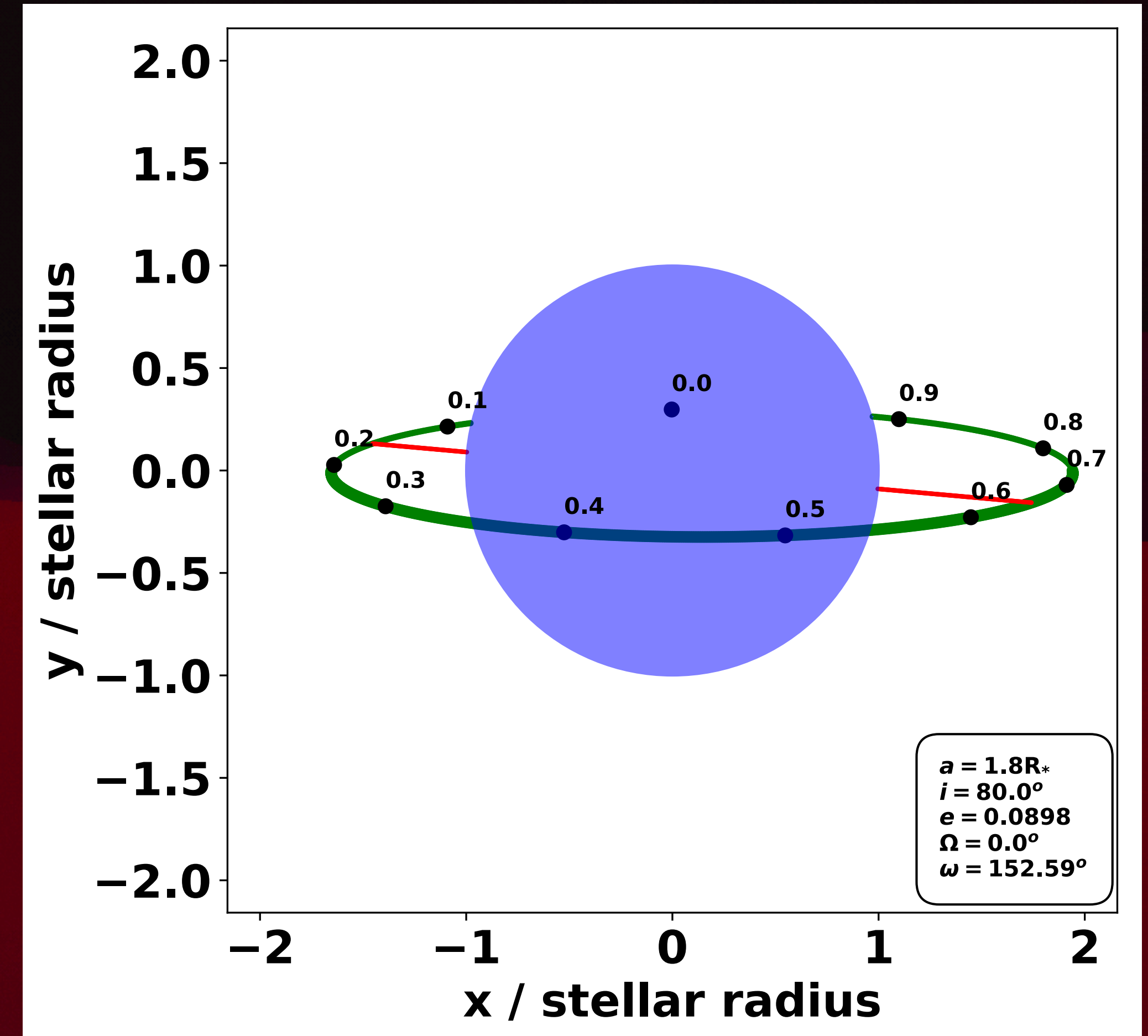
(4) *Rawls et al. (2011)*

(5) *Kreykenbohm et al. (2008)*

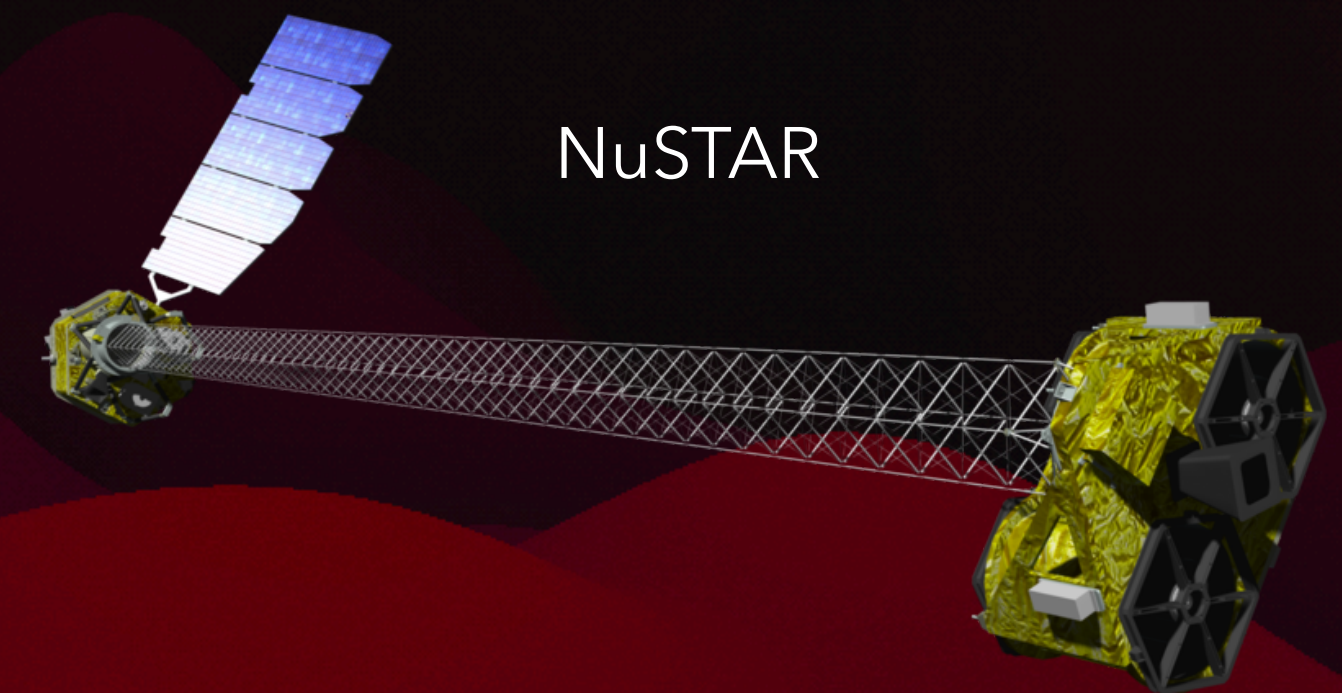
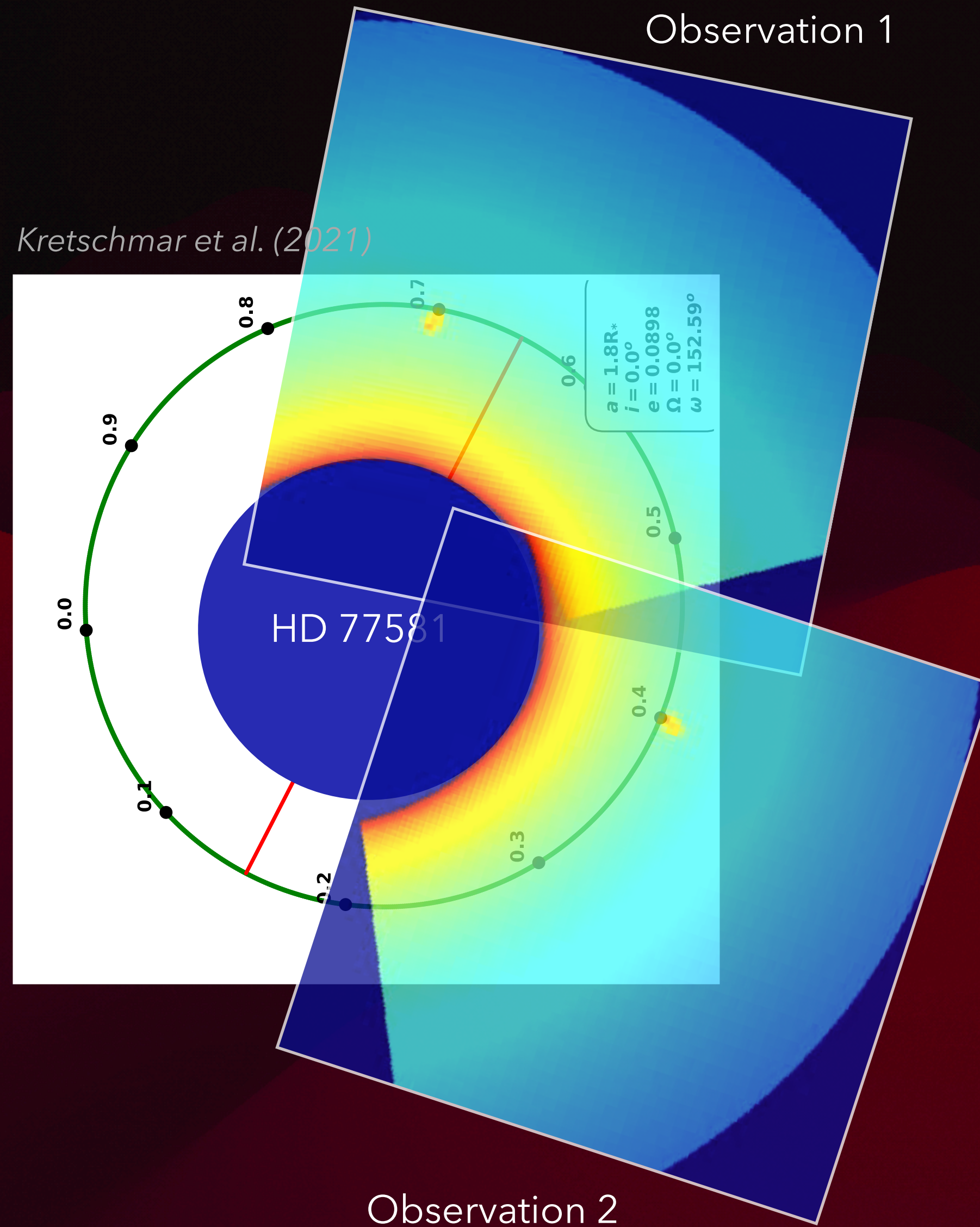
(6) *Bildsten et al. (1997)*

(7) *van Kerkwijk et al. (1995)*

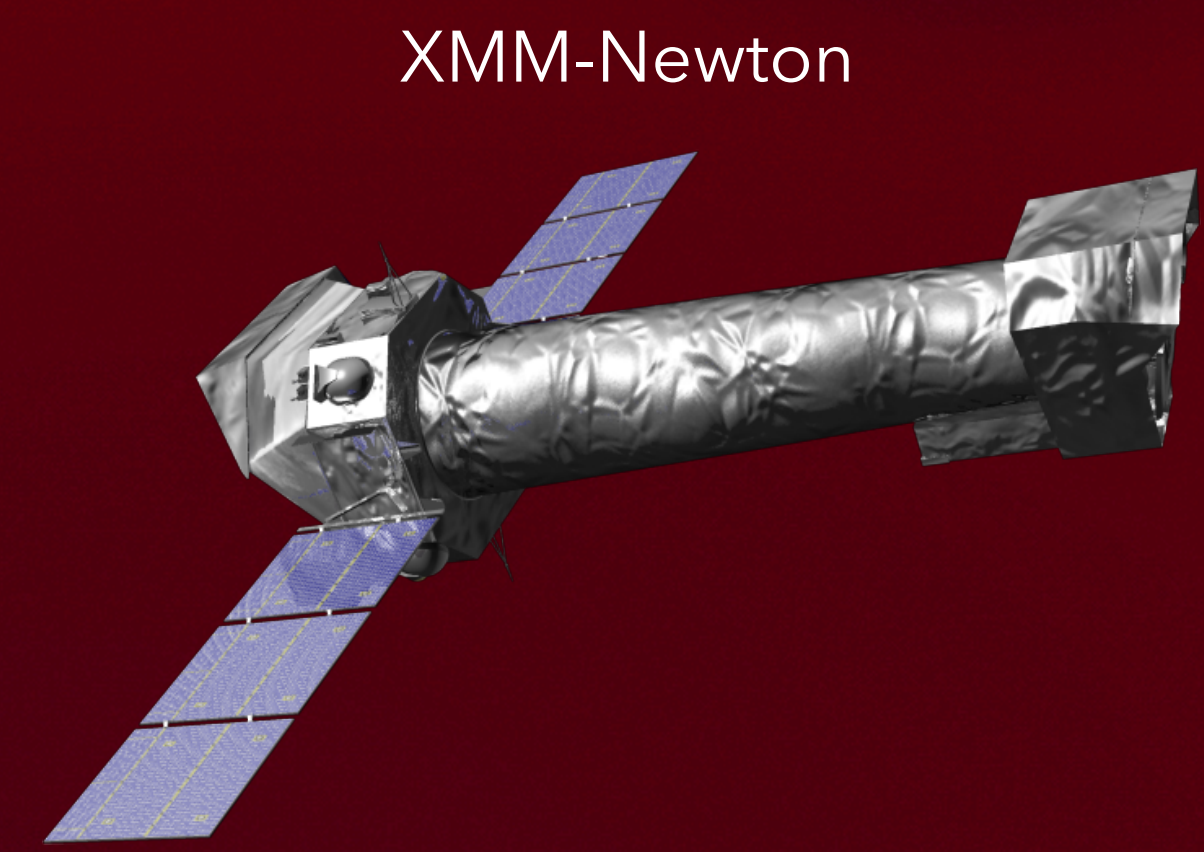
(8) *McClintock et al. (1976)*



An example: Vela X-1

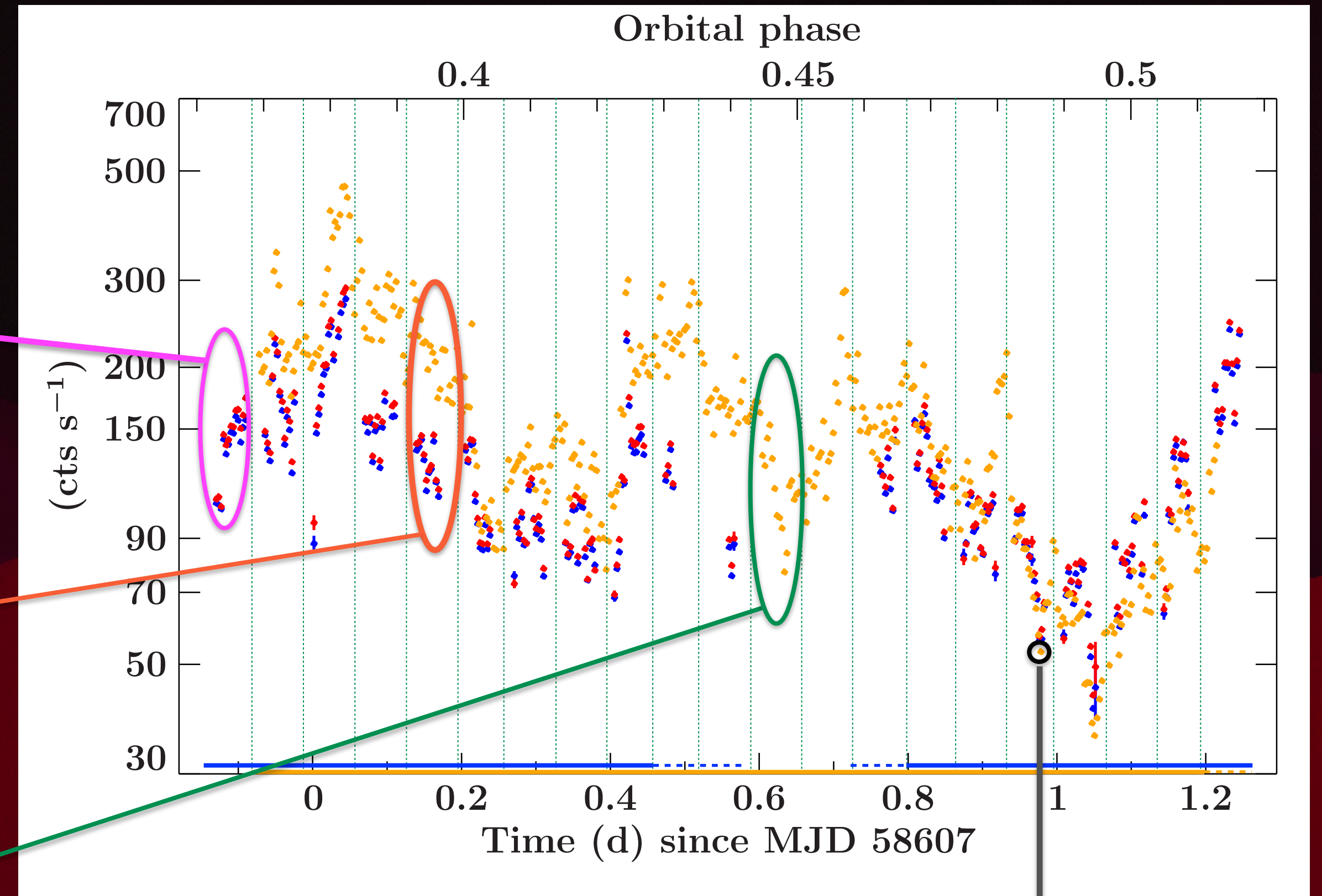
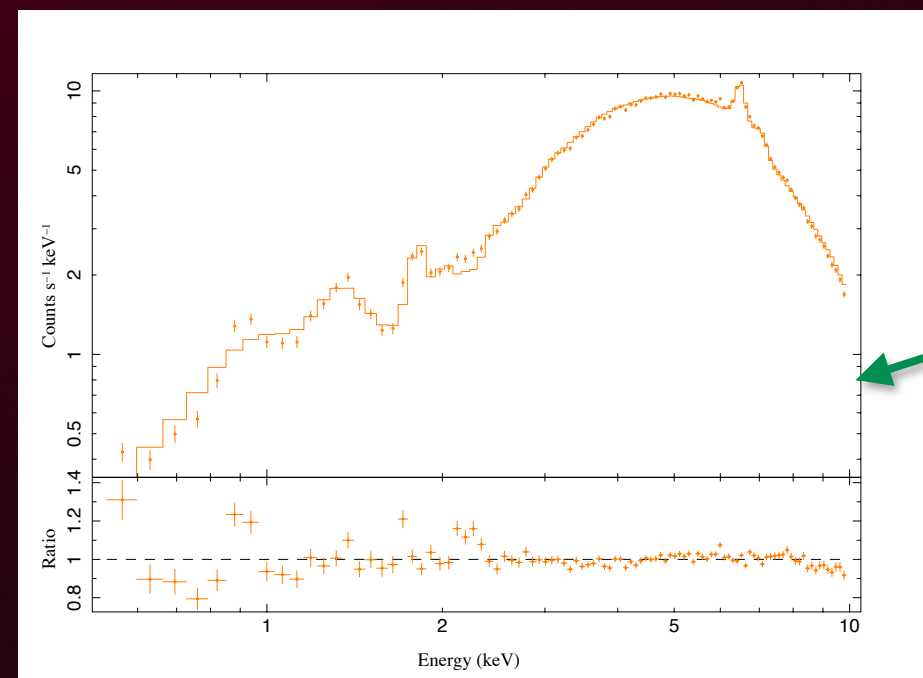
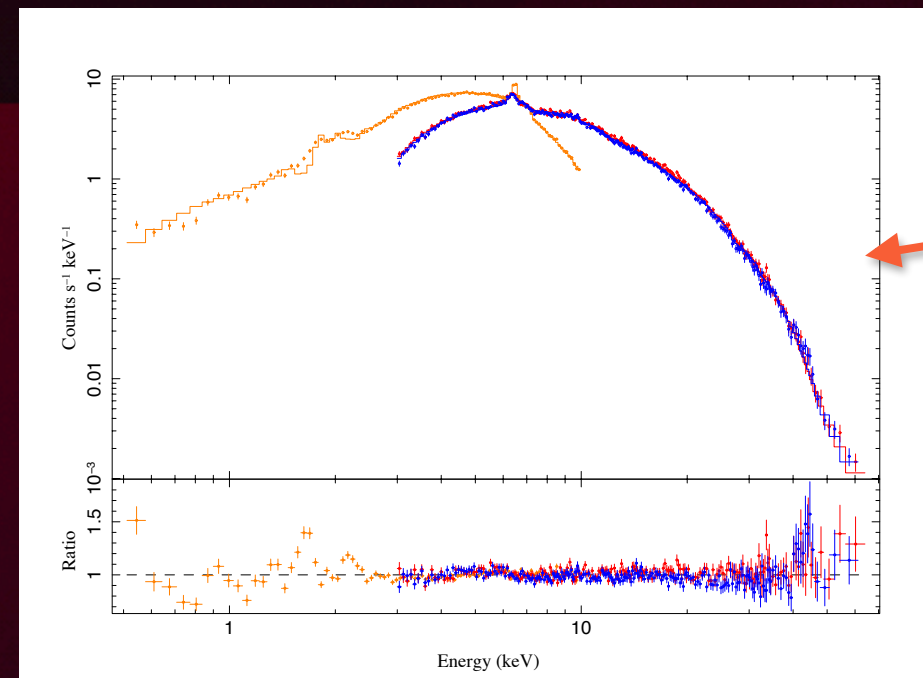
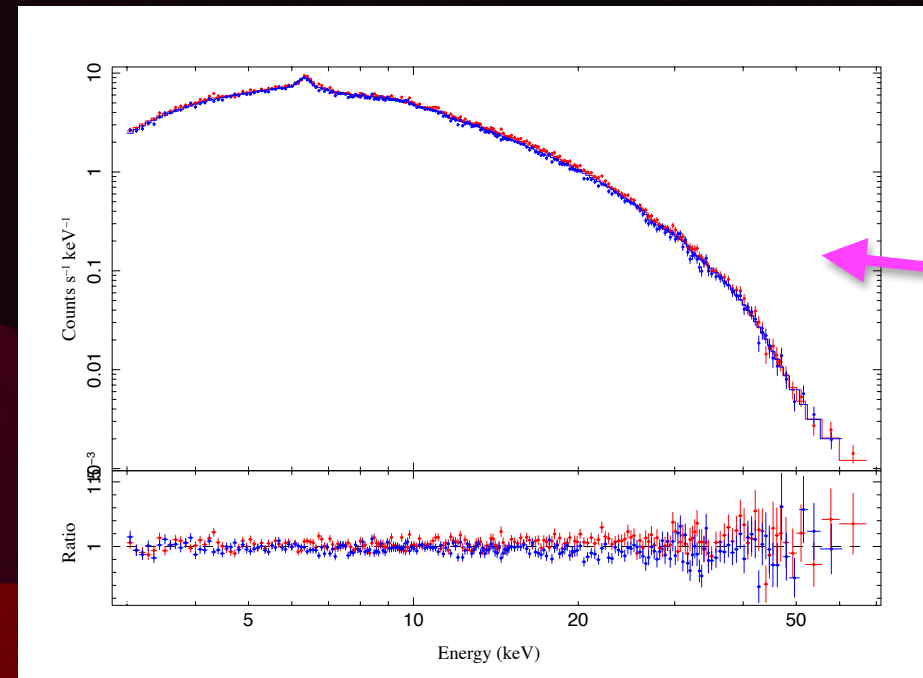


+

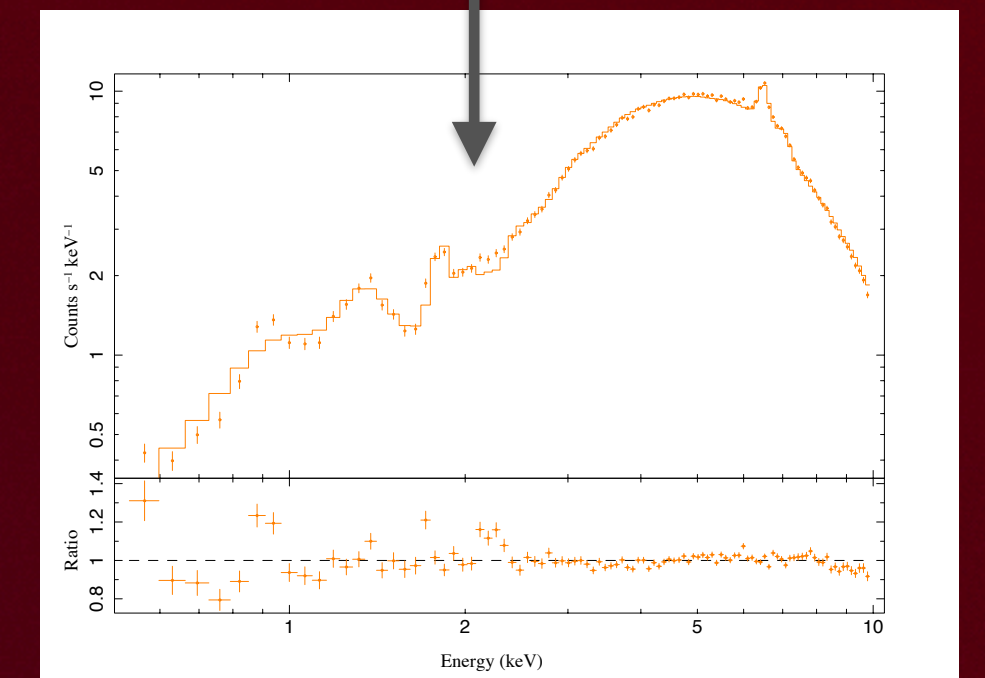


Observation II

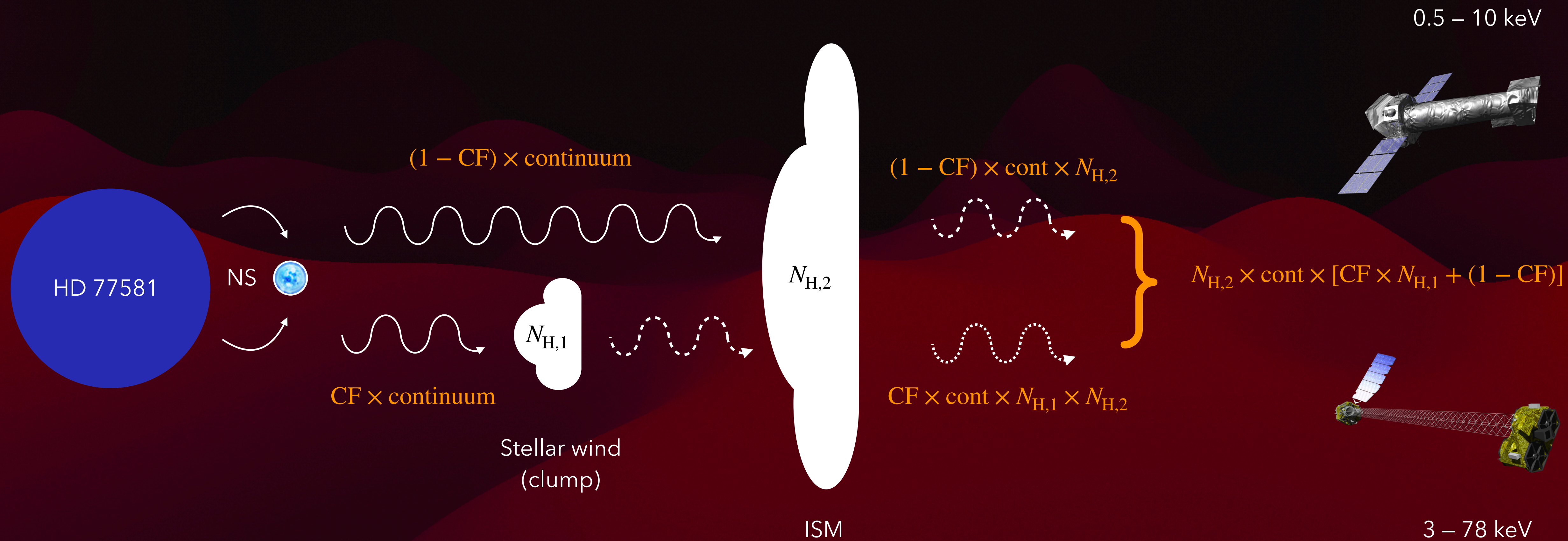
NuSTAR + XMM-Newton



Diez et al. (2023)



Partial Covering Model

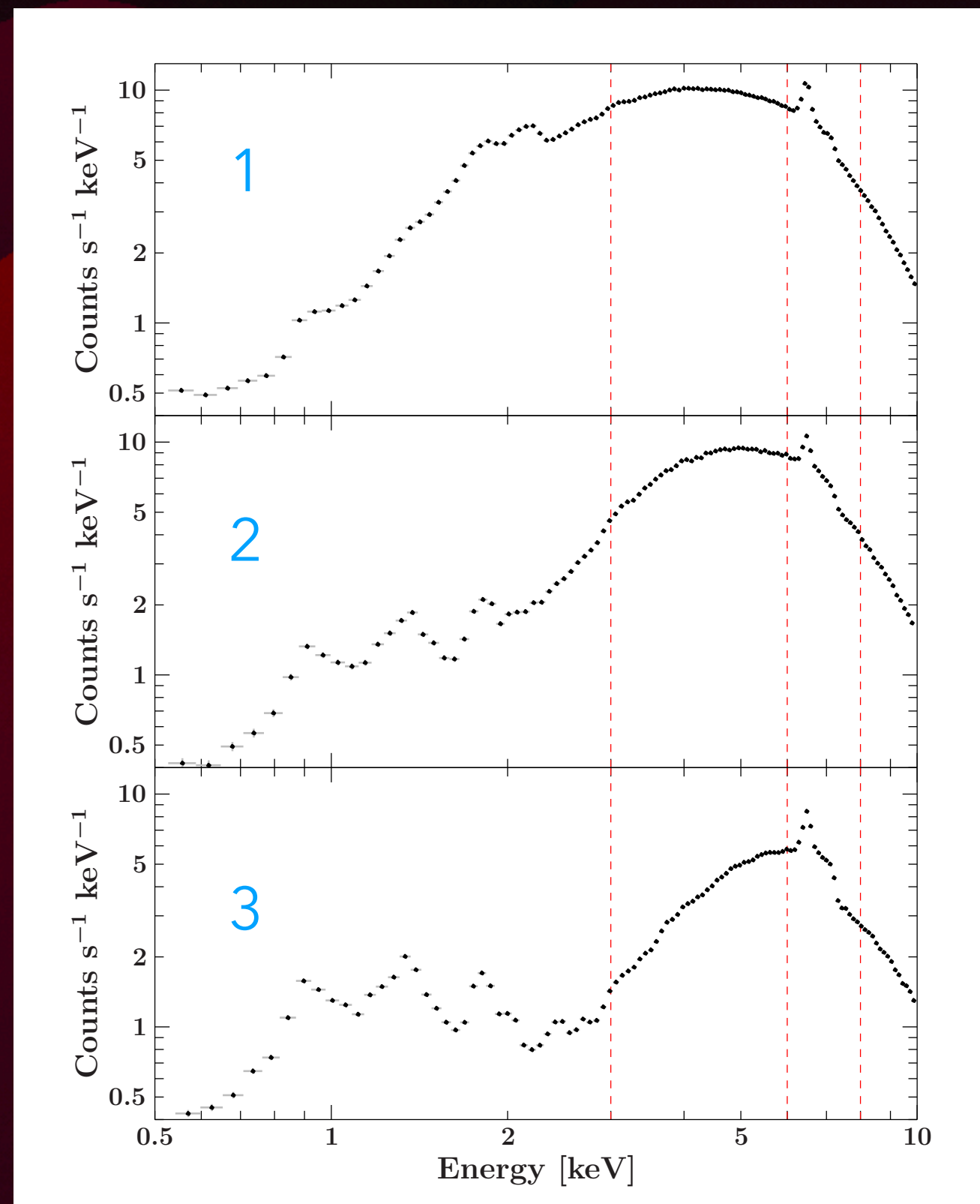


(not to scale)

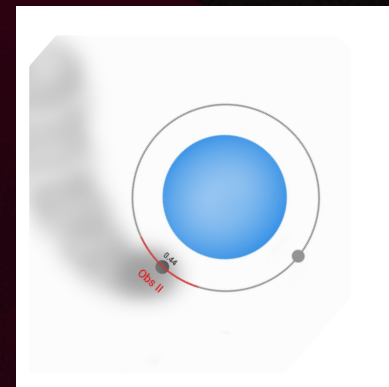
XMM-Newton results (Using constraints obtained with NuSTAR)

Absorption variability + photoionisation of the wind

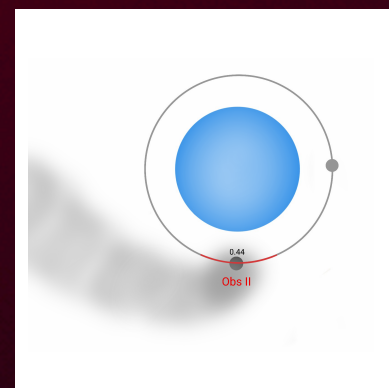
Diez et al. (2023)



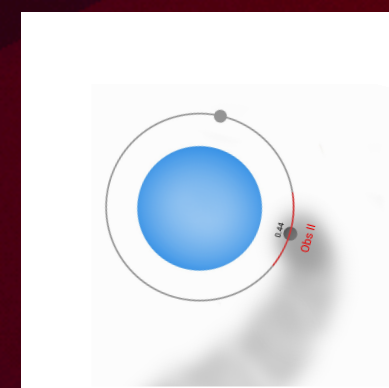
Time



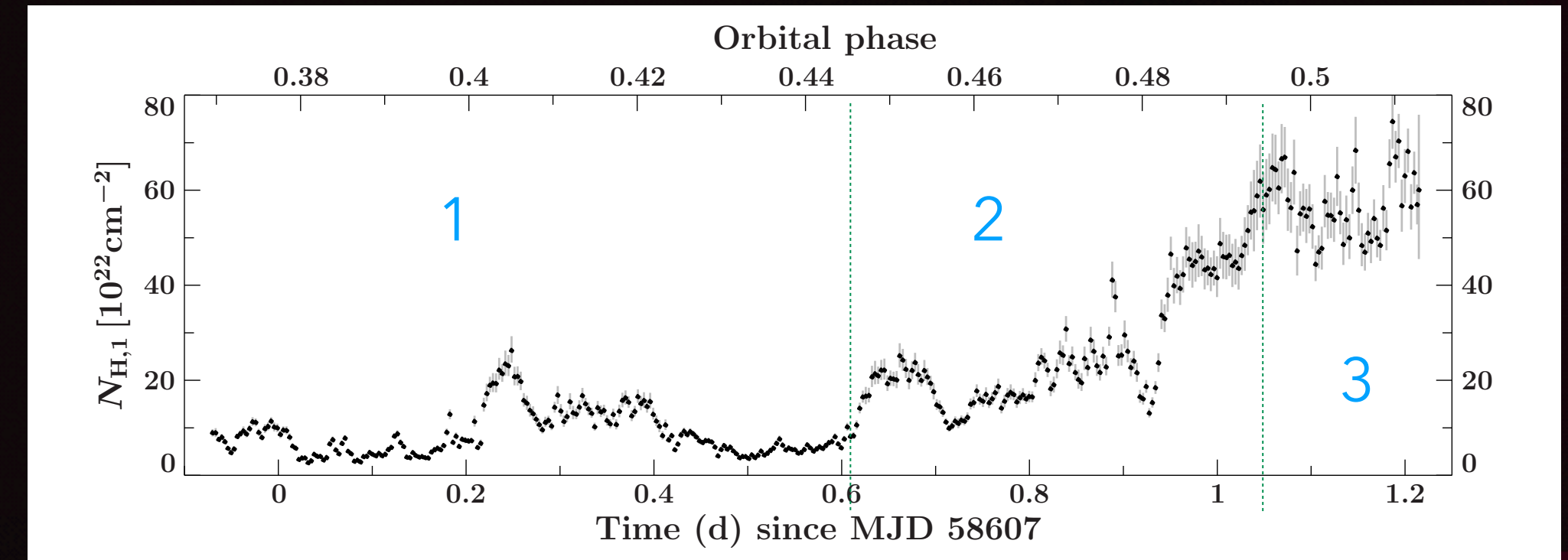
1



2



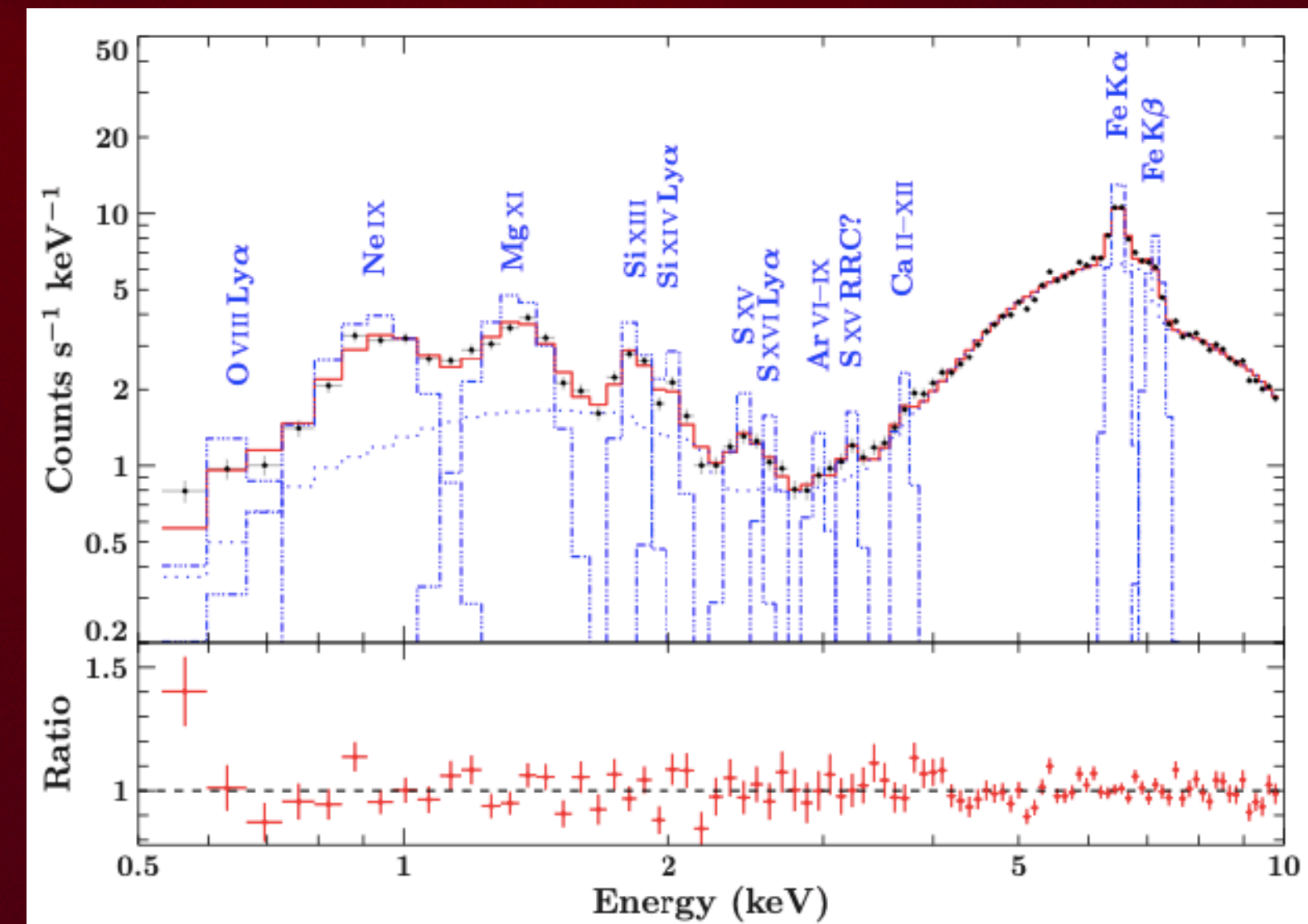
3



► Strong rise of $N_{H,1}$

Diez et al. (2023)

Diez et al. (2023)

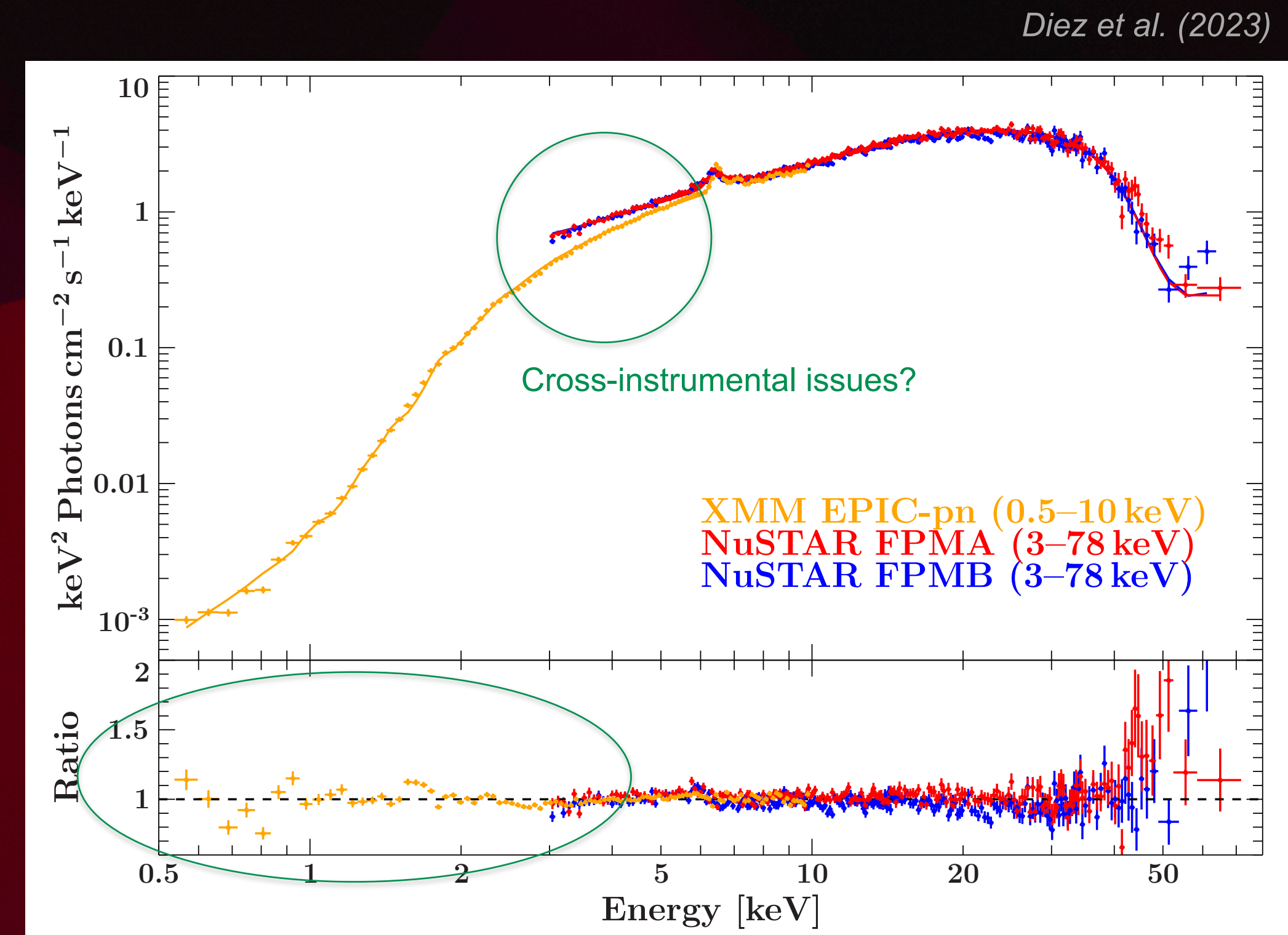


► Photoionisation + lines originate from different regions

Conclusions

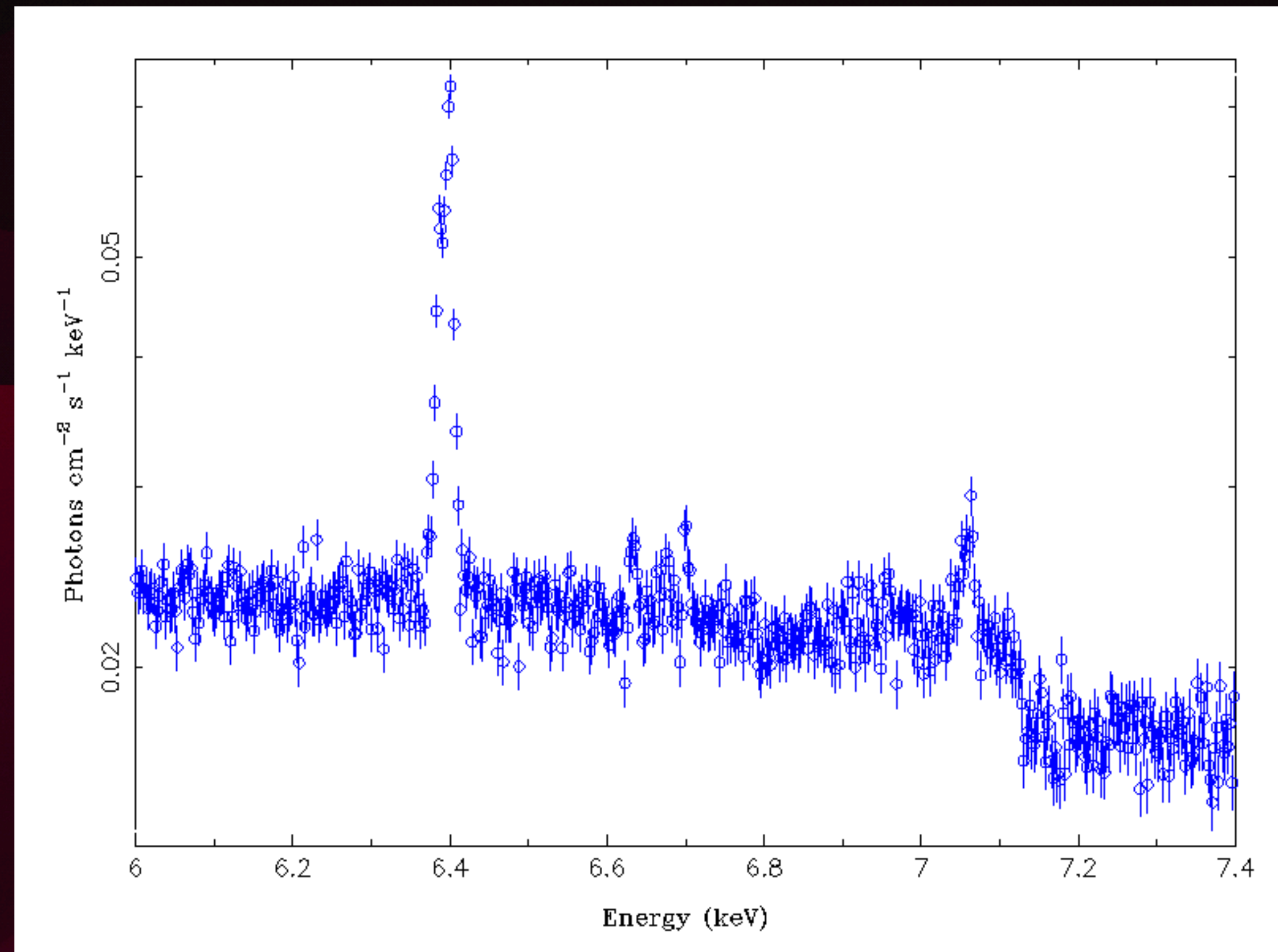
- ▶ Such a detailed onset of the stellar wind in Vela X-1 on a broad X-ray range for the first time
- ▶ Strong photoionisation of the wind
- ▶ Robust model for partially covered sources

- ▶ Cross-calibration issues between XMM-Newton and NuSTAR, seems to be recurrent for piled-up sources



Last update with XRISM

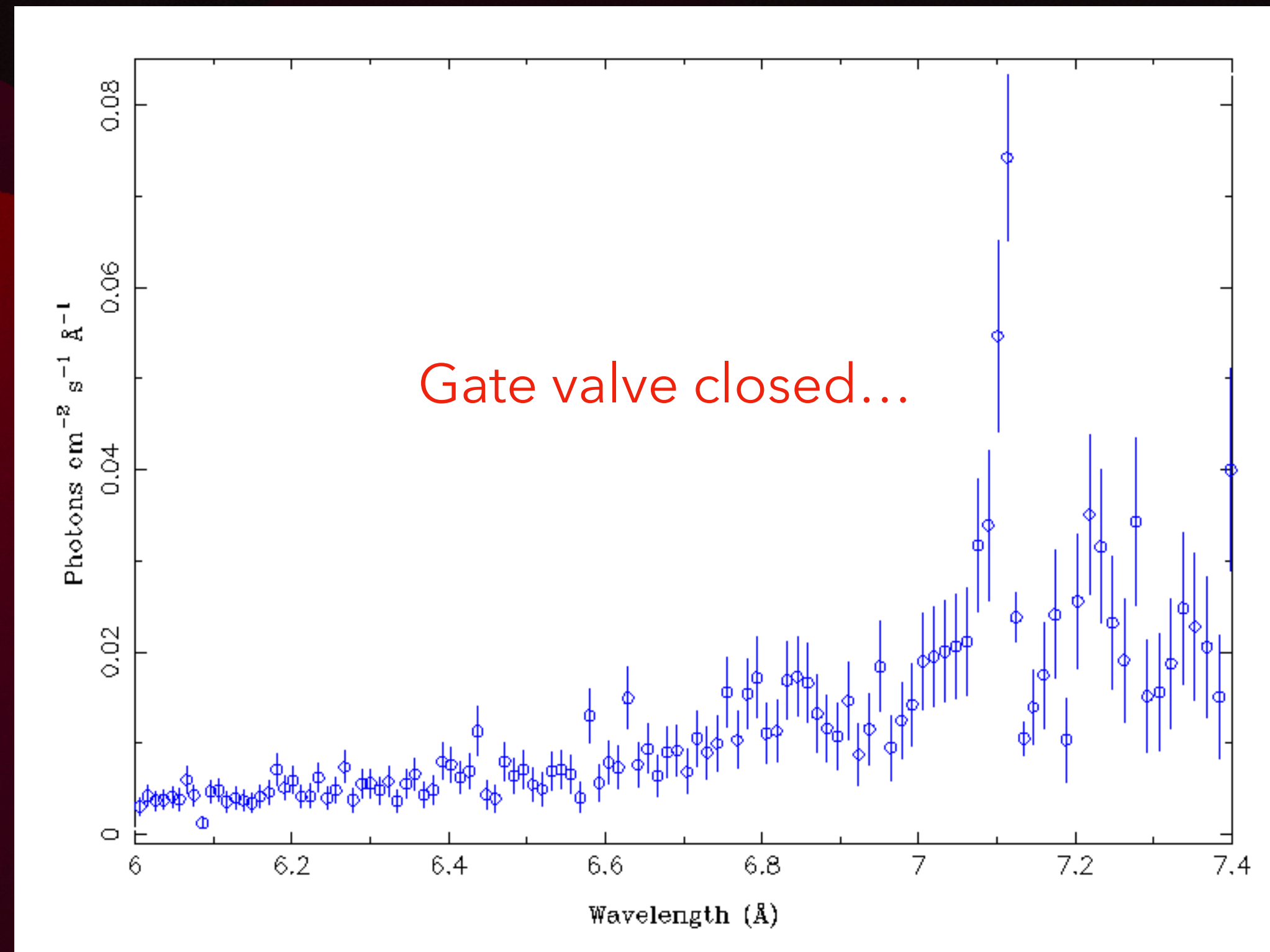
Highly-resolved iron line region



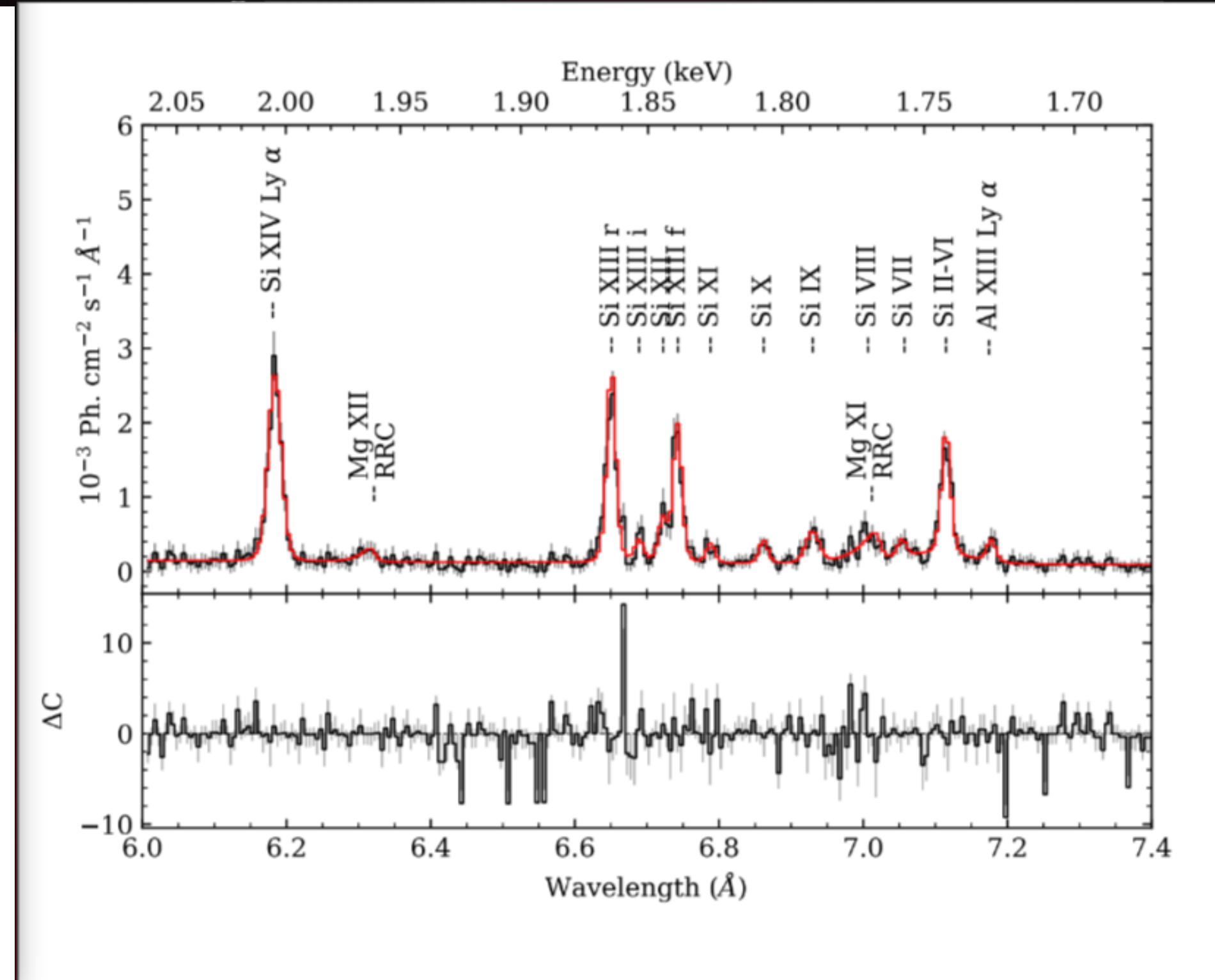
Last update with XRISM

But that's all...

XRISM Resolve



Chandra HETGS

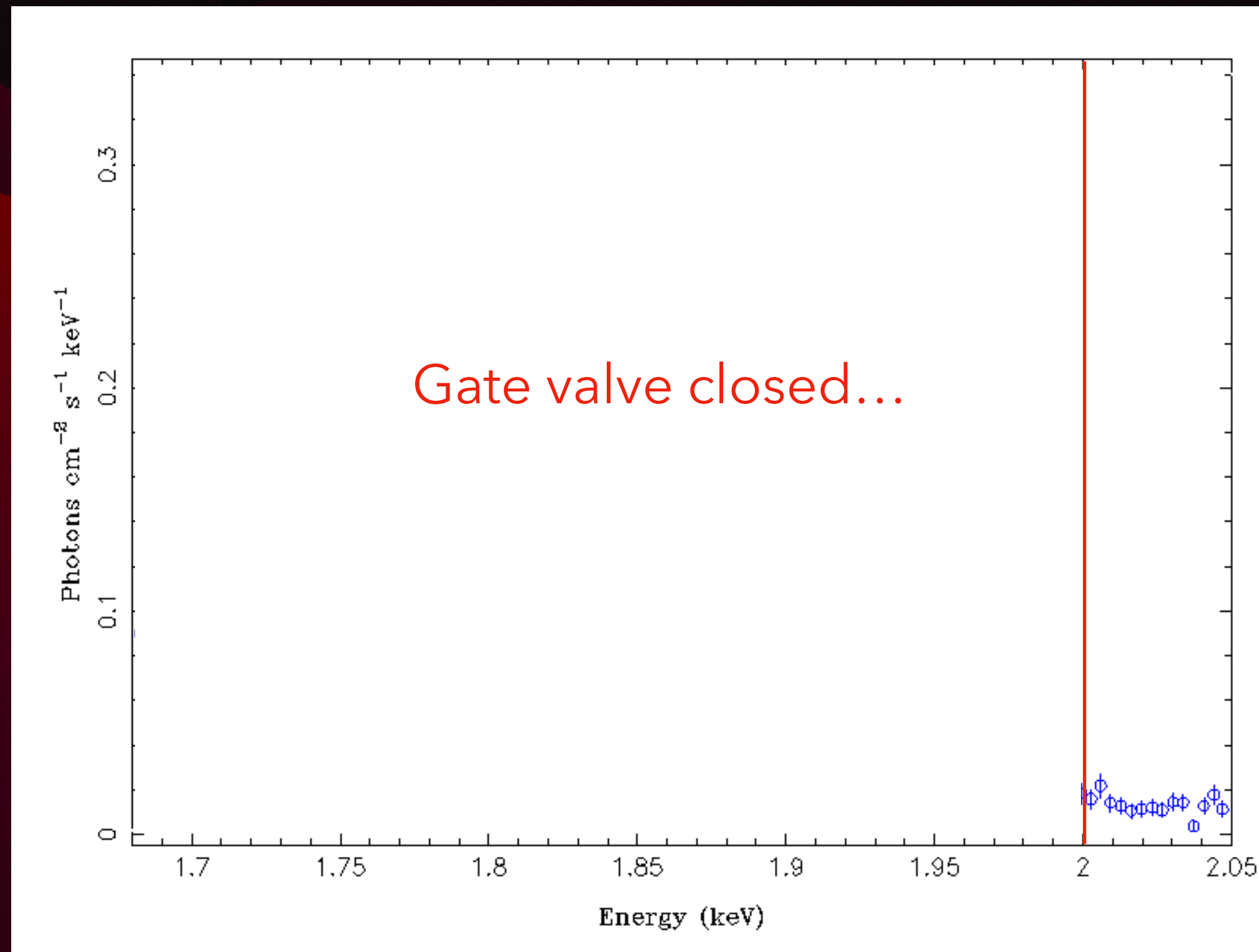


Amato et al. (2020)

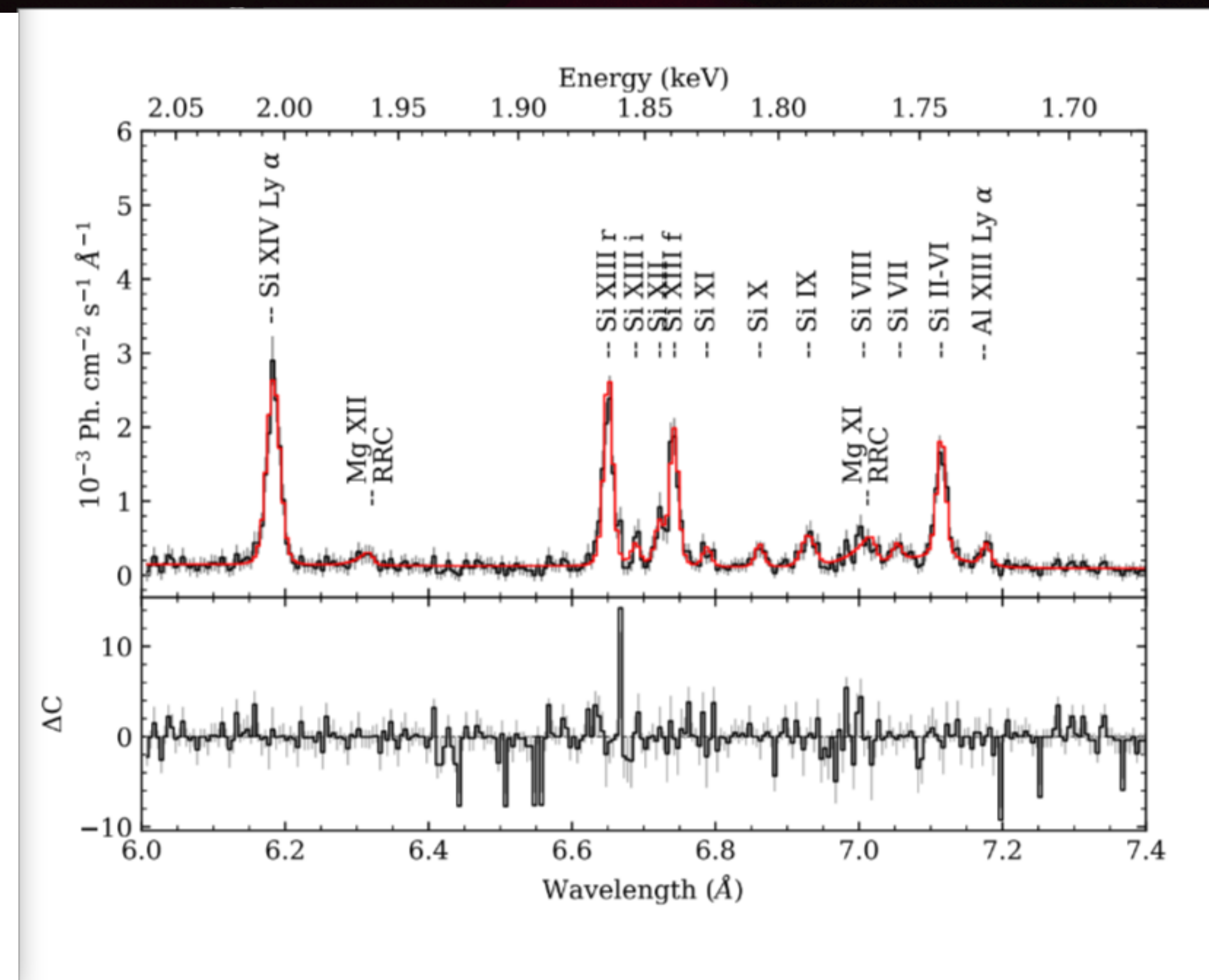
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Amato et al. (2020)